



@CANDELS_team

<http://candels-collaboration.blogspot.com>

Morphological Classification of Galaxies in CANDELS

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+ the CANDELS Collaboration

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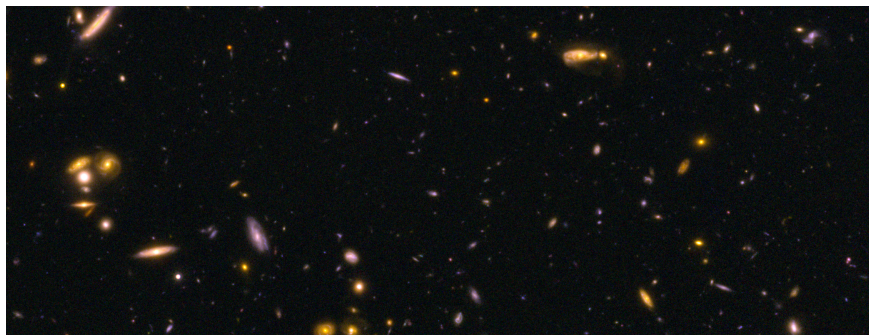
2013 September 25

Galaxy Zoo: Evolutionary Paths in Galaxy Morphology

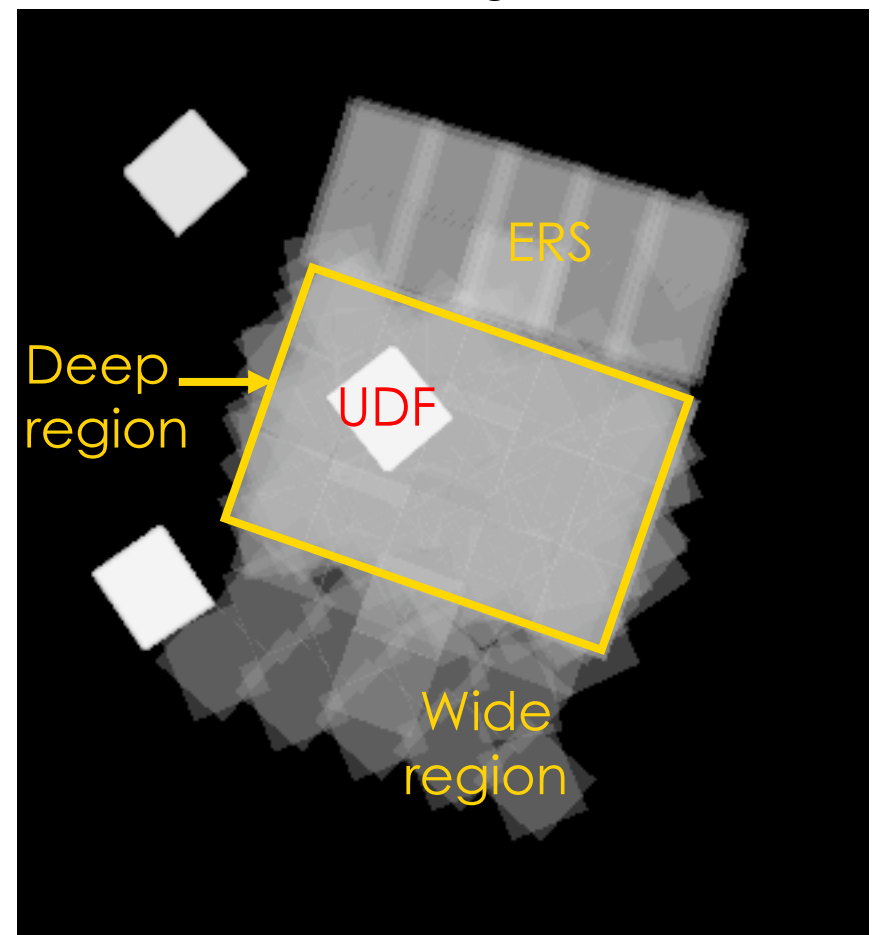
CANDELS



- Largest HST Survey (PIs: S. Faber & H. Ferguson)
- NIR imaging of 5 of the most commonly studied deep fields
- Deep and wide areas
- Imaging is now complete!
- Work is ongoing



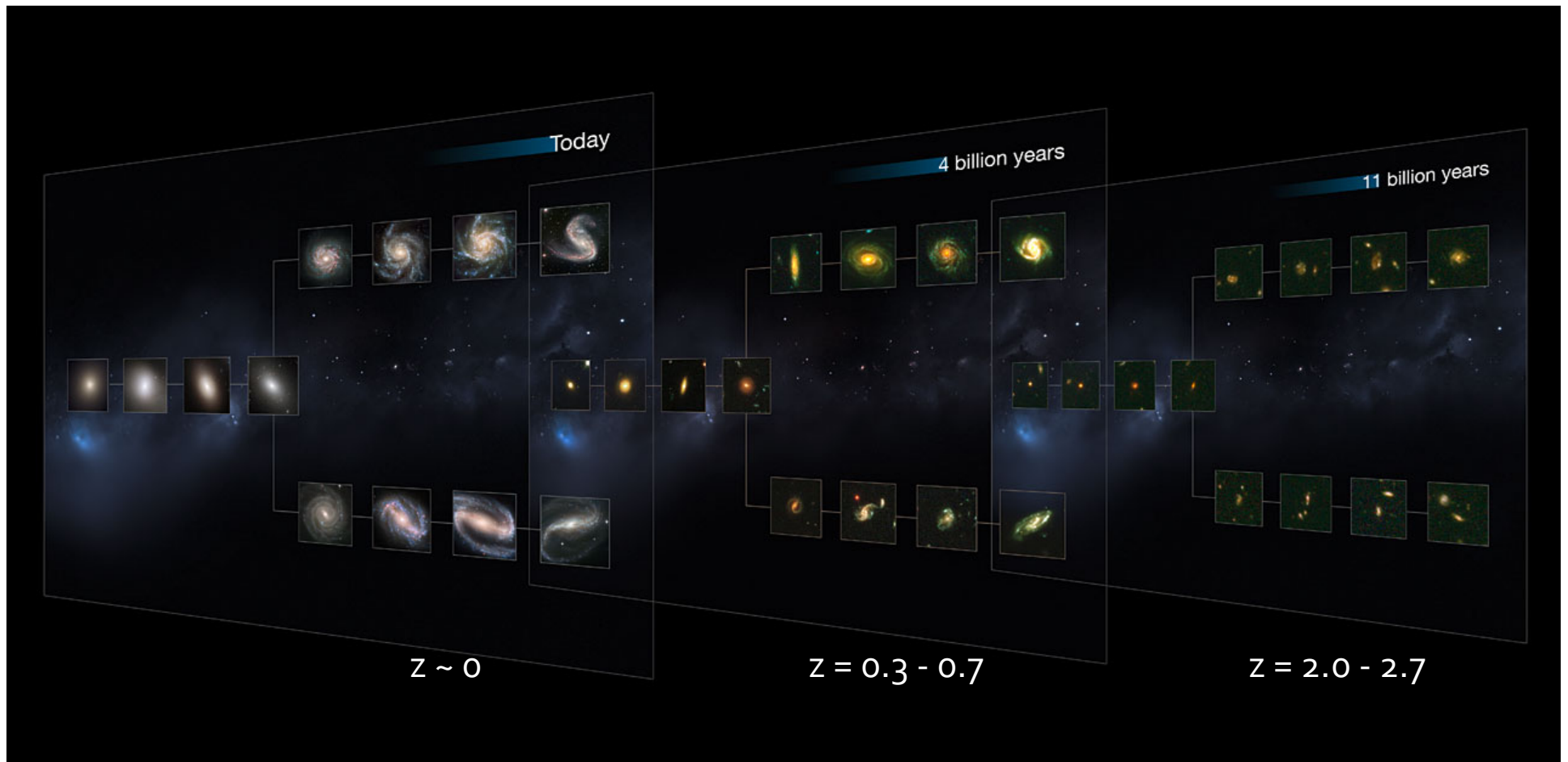
J and H band Coverage of GOODS-S



Galaxy Morphology at High- z

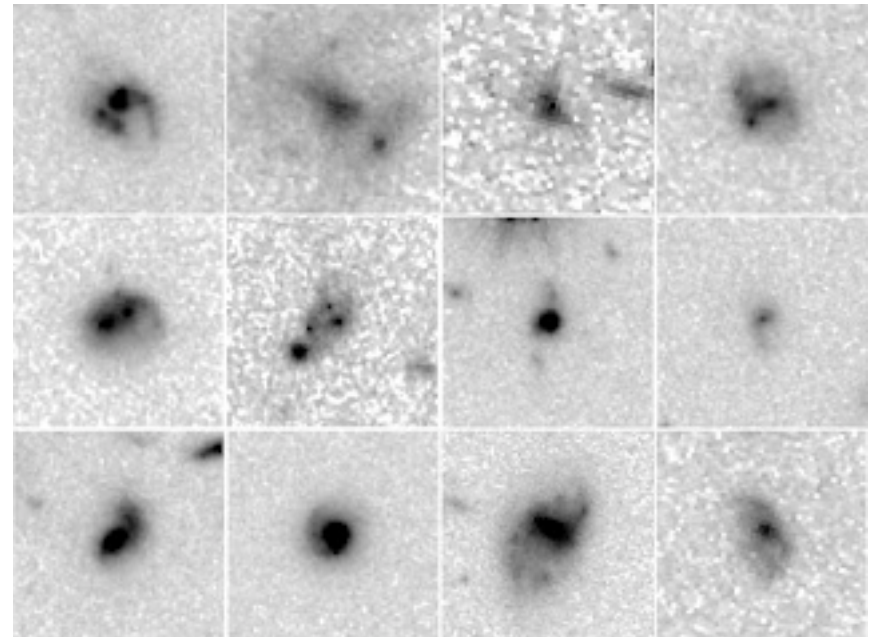
- For galaxies at $z \sim 2$, deep NIR imaging provides rest-frame optical structure
- With CANDELS WFC3, can look at structure for large samples at $z \sim 2$ for the first time
- Visual Morphology
- Quantitative Morphology (see talk by Jen Lotz)
 - GALFIT: Sersic index, B/D, size, etc.
 - Non-parametric measures: Gini, M20, CAS, MID, etc.
 - Use Visual Morphology to calibrate these at high- z

When did the Hubble Sequence Form?



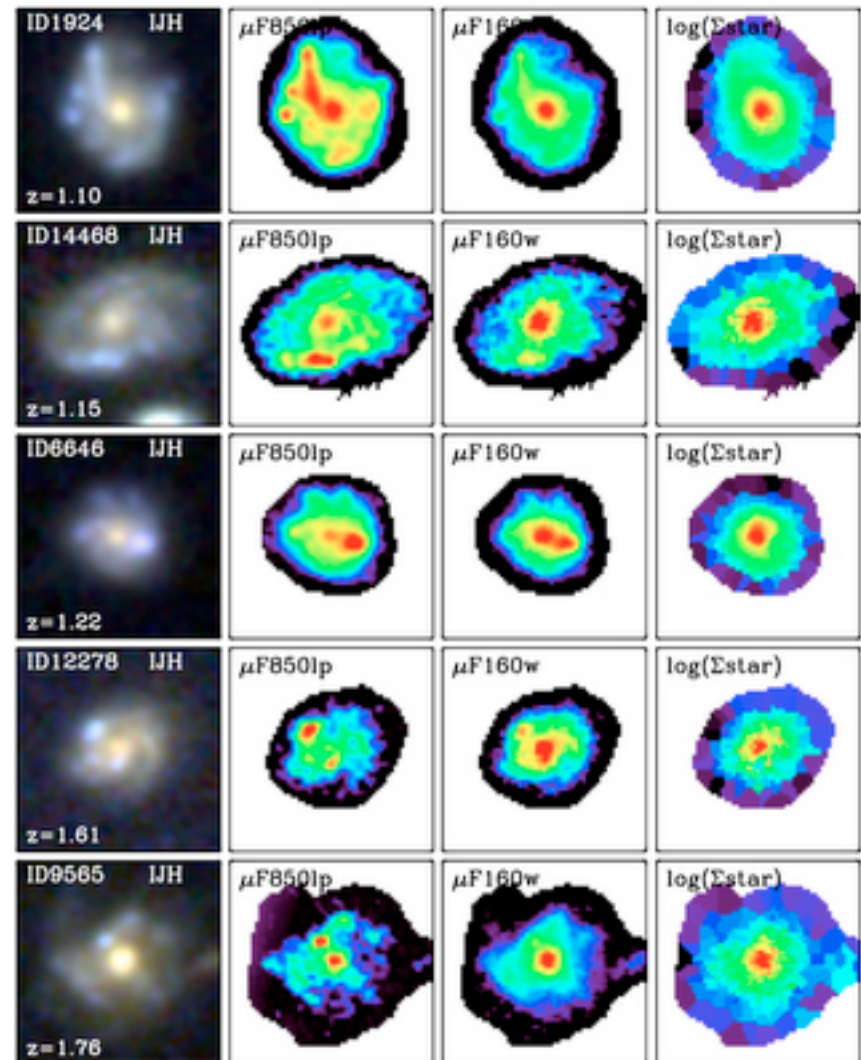
What is the Role of Galaxy Mergers?

- In increasing a galaxy's SFR
- In contributing to the cosmic star formation history
- In fueling AGN
- In the morphological transformation of galaxies
 - Forming spheroids
- Hints at the changing role of mergers in the recent literature but many open questions remain!



How Important are Clumps?

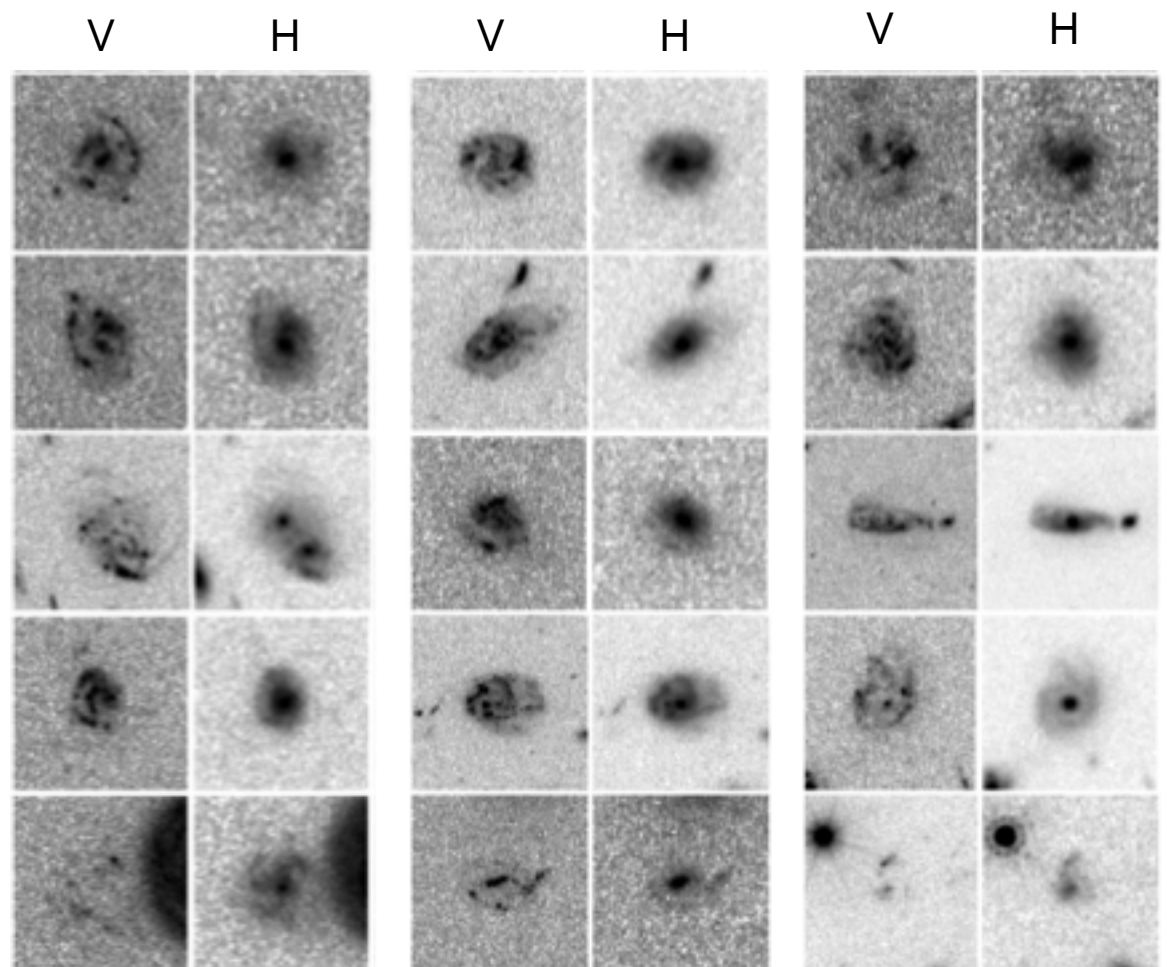
- Disk instabilities?
- Signatures of mergers?
- More frequent in higher gas fraction systems?
- Do these clumps become part of the bulge?



Wuyts et al. 2012

Morphological K-Corrections

- Some objects are very different in the optical and NIR
- Many clumpy irregular systems look regular in the IR
- Bulges are more prominent in the IR



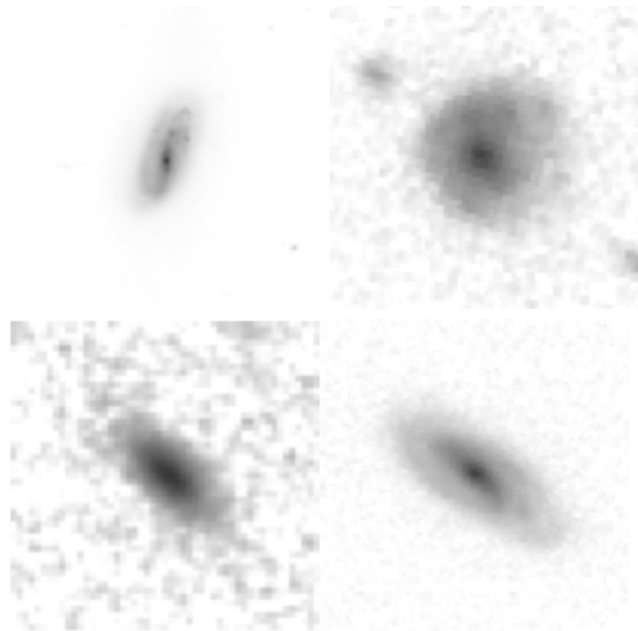
CANDELS Visual Classification Effort

- Classifying all CANDELS Galaxies to $H < 24.5$
 - 3 fields complete (GOODS-S, UDS, COSMOS)
 - 2 remaining (GOODS-N underway)
 - Total of ~50,000 galaxies by the end of survey
 - Classifications at multiple depths in deep area
 - Primarily in H-band, use other bands to inform
- Multiple classifiers for comparison and statistical analysis
 - ~3-5 people per object
 - ~65 classifiers in total
- Catalogs to be made public (GOODS-S with Kartaltepe et al., in prep)

Classification Scheme

- Two levels
 - Main Morphological Class
 - NOT mutually exclusive! – can choose more than one

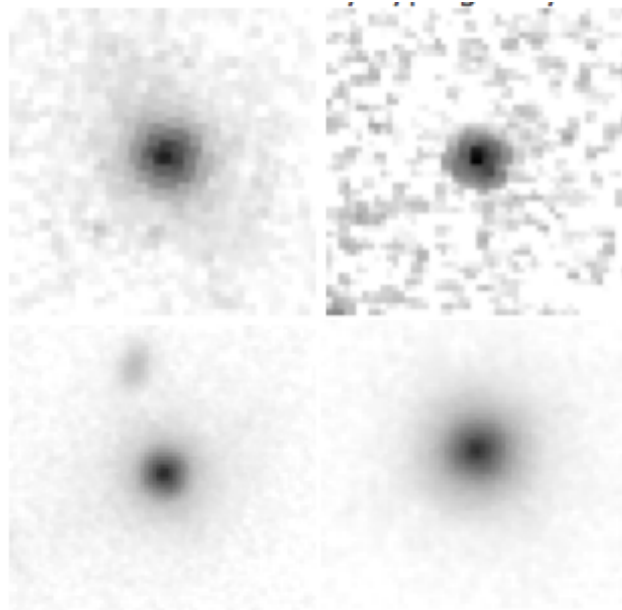
Disks



Classification Scheme

- Two levels
 - Main Morphological Class
 - NOT mutually exclusive! – can choose more than one

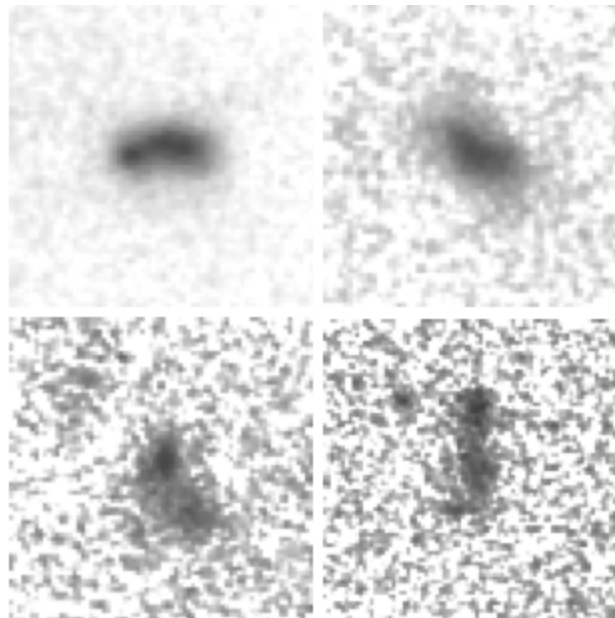
Spheroids



Classification Scheme

- Two levels
 - Main Morphological Class
 - NOT mutually exclusive! – can choose more than one

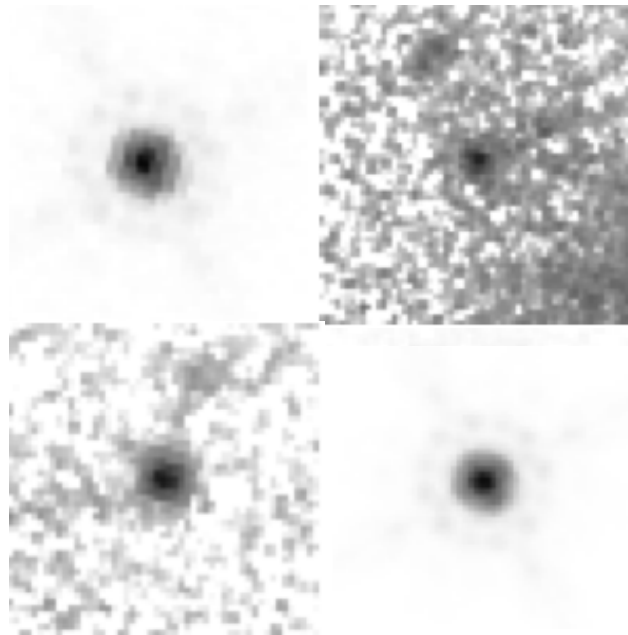
Irregular



Classification Scheme

- Two levels
 - Main Morphological Class
 - NOT mutually exclusive! – can choose more than one

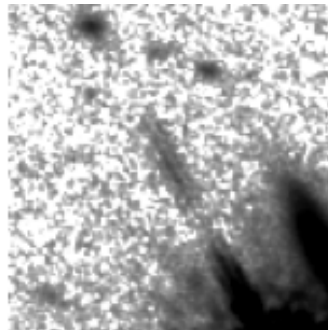
Compact/
Unresolved



Classification Scheme

- Two levels
 - Main Morphological Class
 - NOT mutually exclusive! – can choose more than one

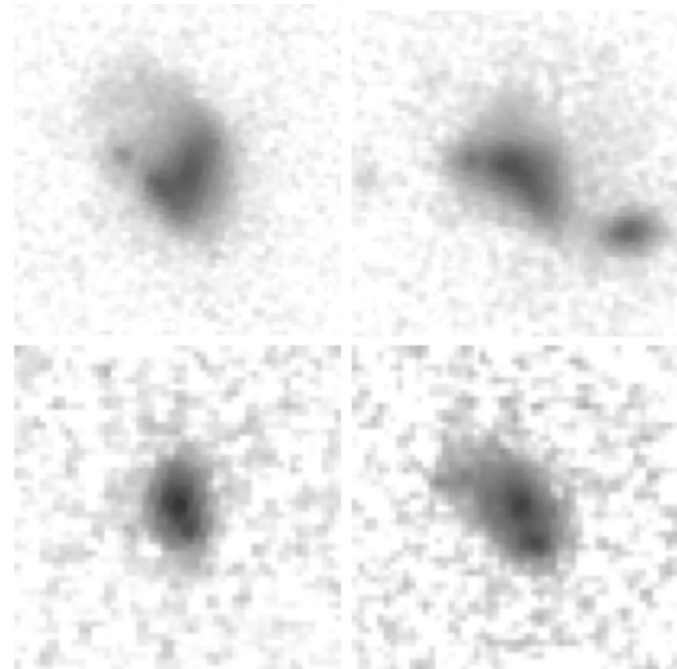
Unclassifiable



Classification Scheme

- Two levels
 - Main Morphological Class (disk, spheroid, irregular, point source, unclassifiable)
 - NOT mutually exclusive! – can
 - Interaction Class
 - Only one choice (or none)

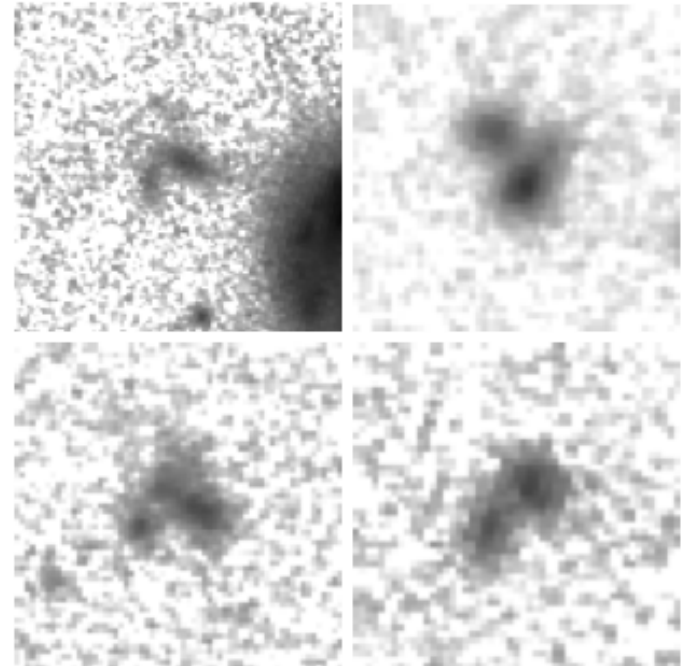
Mergers



Classification Scheme

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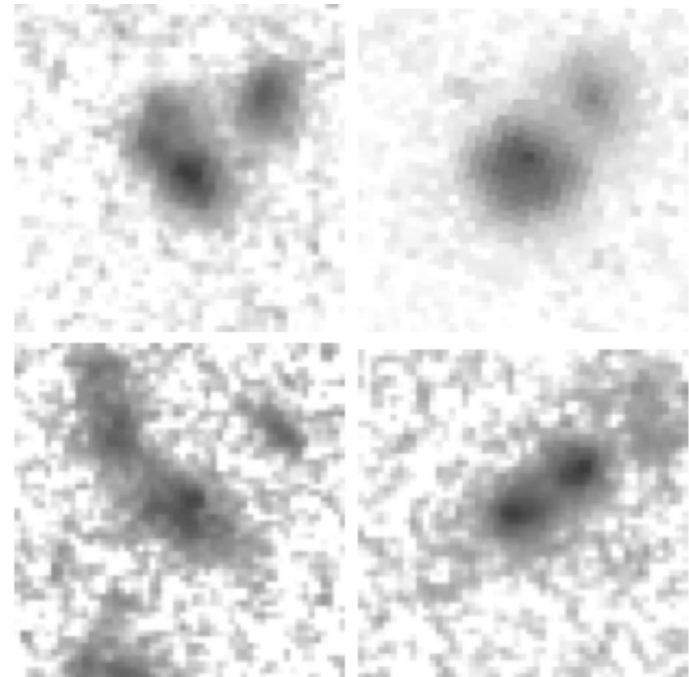
Interaction within
Segmentation Map



Classification Scheme

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 - NOT mutually exclusive! – can
 - Interaction Class
 - Only one choice (or none)

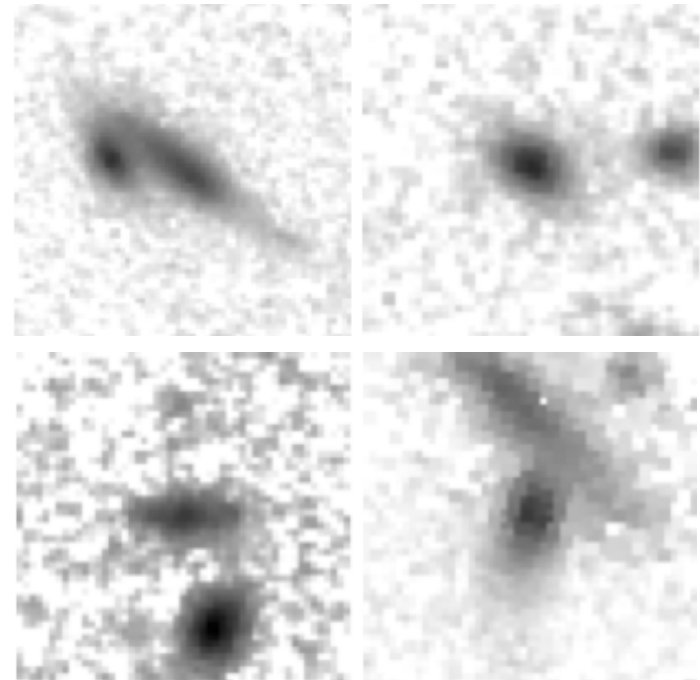
Interaction Beyond
Segmentation Map



Classification Scheme

- Two levels
 - Main Morphological Class (disk, spheroid, irregular, point source, unclassifiable)
 - NOT mutually exclusive! – can
 - Interaction Class
 - Only one choice (or none)

Non-interacting
Companion



Classification Scheme

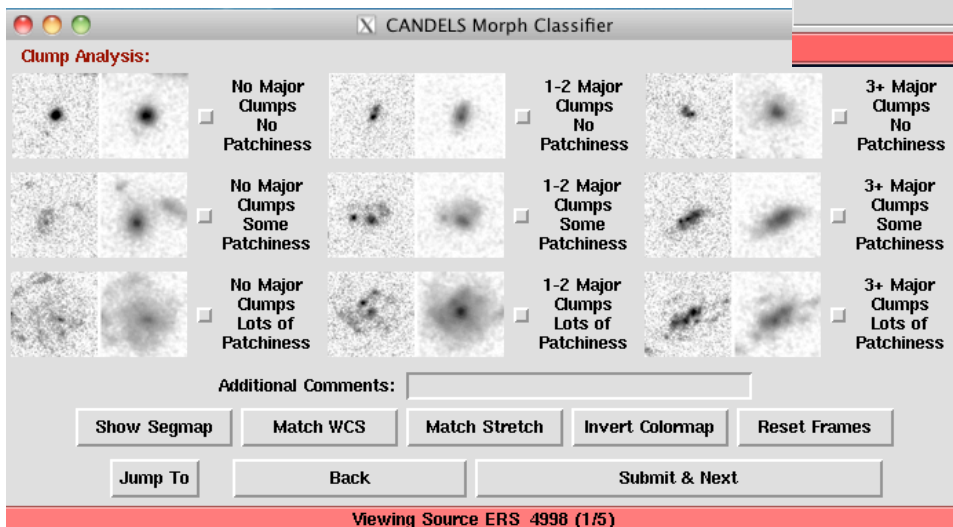
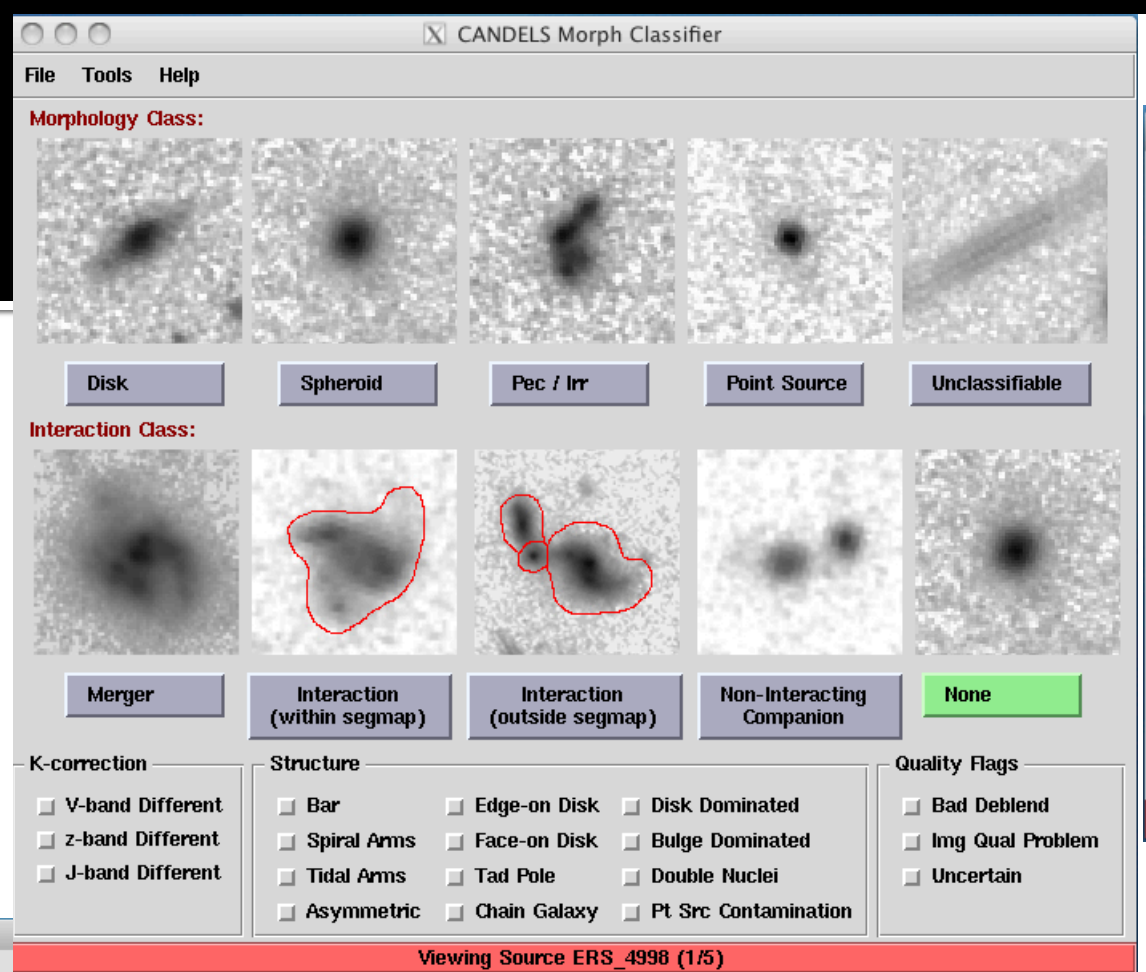
- Two levels
 - Main Morphological Class (disk, spheroid, irregular, point source, unclassifiable)
 - NOT mutually exclusive! – can choose more than one
 - Interaction Class
 - Only one choice (or none)
- Structure flags
 - Tidal arms, double nuclei, asymmetric, spiral arms, bar, point source contamination, edge-on/face-on disk, bulge/disk dominated, tadpole, chain

Classification Scheme

- Two levels
 - Main Morphological Class (disk, spheroid, irregular, point source, unclassifiable)
 - NOT mutually exclusive! – can choose more than one
 - Interaction Class
 - Only one choice (or none)
- Structure flags
- K-correction flags
- Quality flags
- Clumpiness flags

GUIs

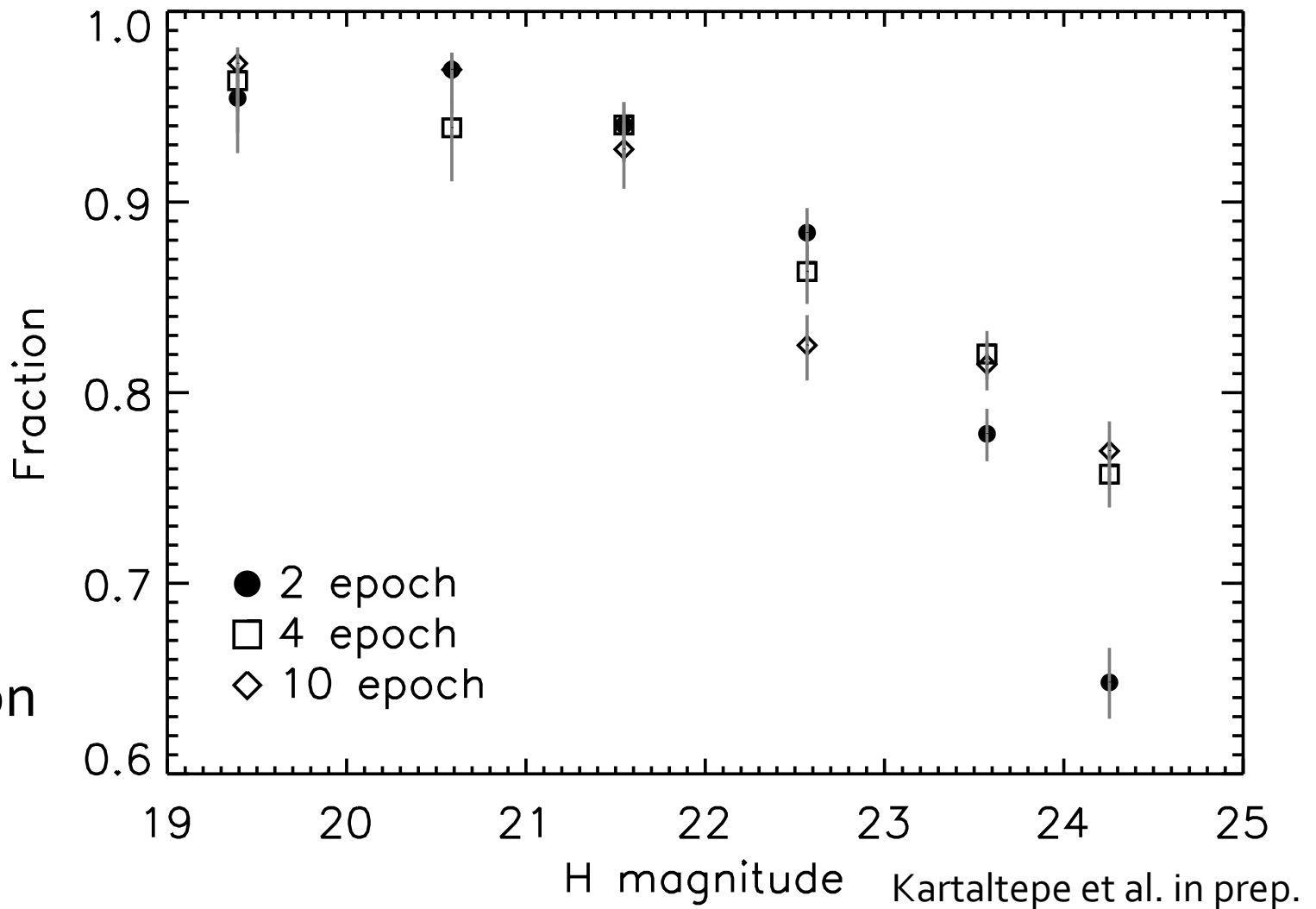
- Web GUI
 - Designed by Mark Mozena
- ds9/Perl GUI
 - Designed by Dale Kocevski



Agreement as a fn of Magnitude

Fraction of galaxies where $> 3/5$ classifiers agree

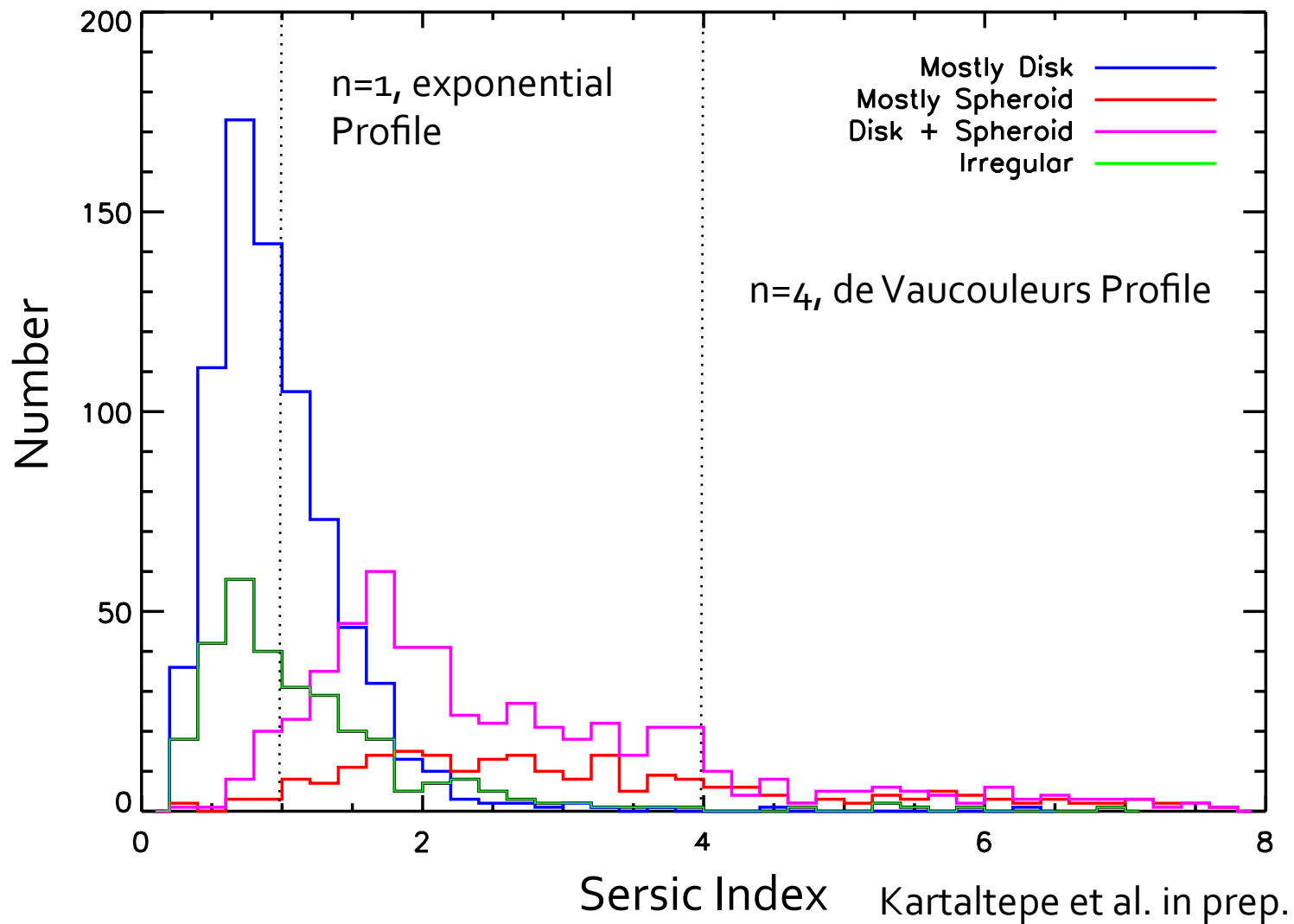
This is also dependent on the classification itself



Lessons Learned So Far

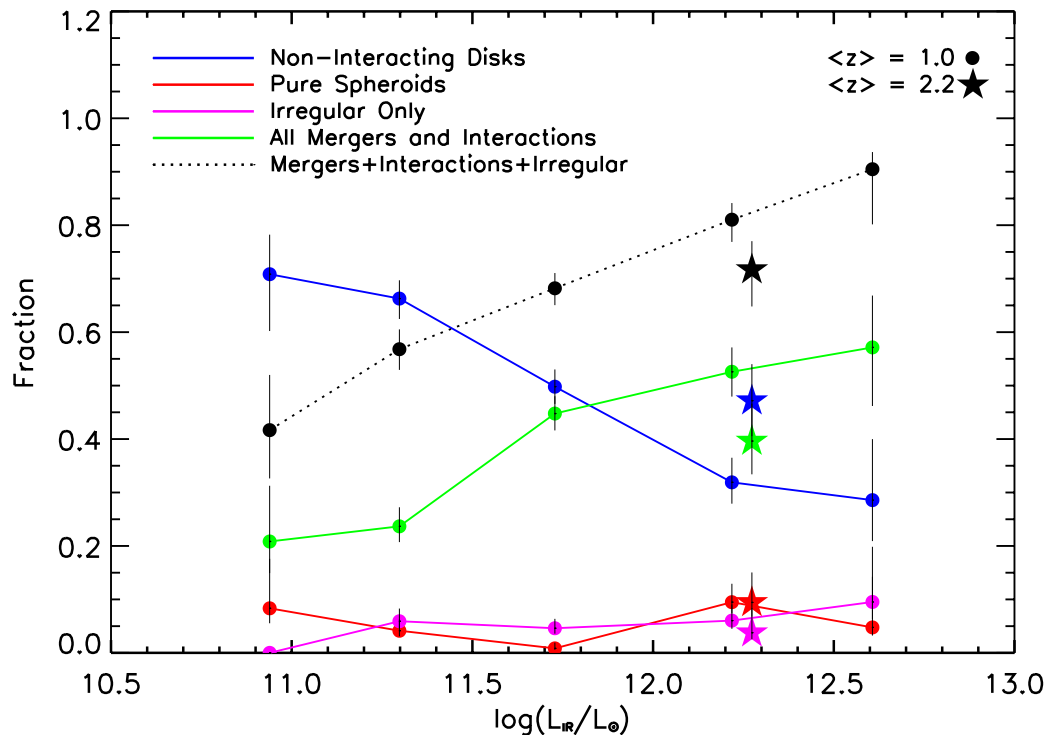
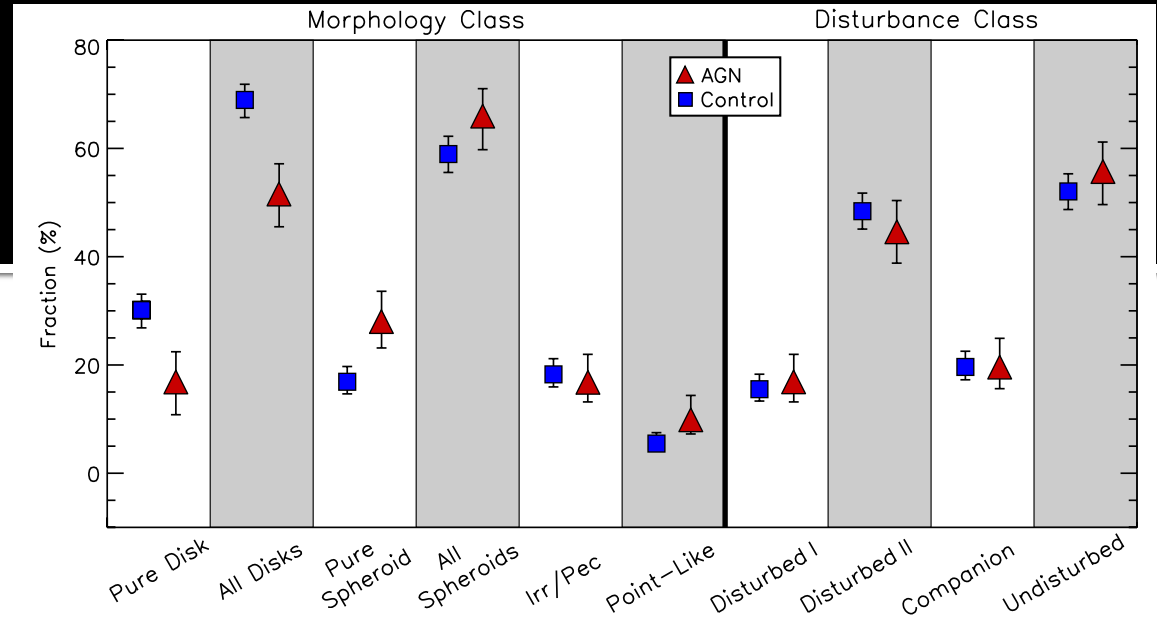
- Different depths
 - Agreement between classifications at 2, 4, and 10 epoch depth is magnitude dependent
 - Small difference at $H > 24$ for disks, $H > 23$ for spheroids
- Relative agreement
 - Agreement among classifiers depends on magnitude
 - Highest agreement for disks and spheroids
 - Lowest agreement for irregulars
 - Generally more complex morphologies
- Most difficult cases

Comparison to Sersic Index



Science Results

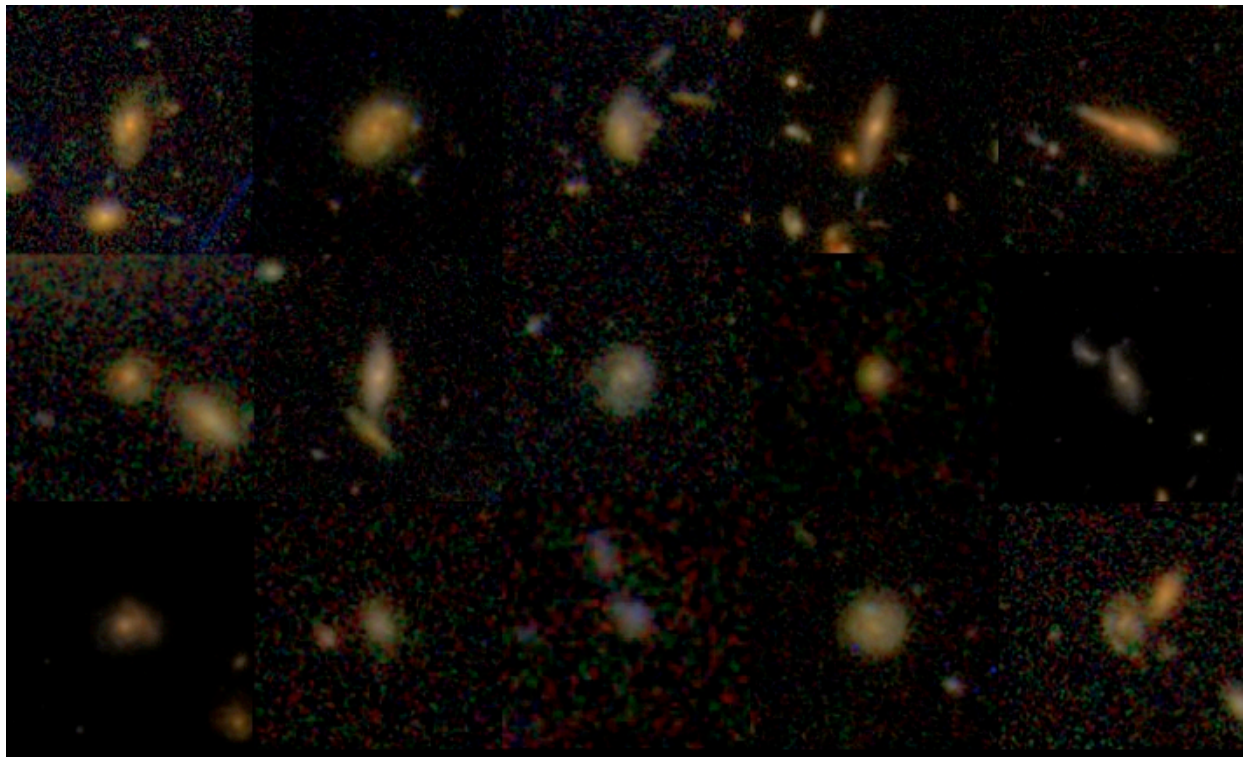
■ Kocevski et al. 2012
Morphology of $z \sim 2$
Moderate Luminosity
X-ray AGN



■ Kartaltepe et al. 2012
Morphology of $z \sim 2$
ULIRGs

CANDELS and Galaxy Zoo

- Postage stamps for all galaxies with $H < 25.5$
 - Color jpegs (IJH – consistent for all fields)
- GOODS-S, UDS, COSMOS
 - Two depths in deep area for comparison
 - Classifications completed
 - Other fields to be added soon
- Plans to add CANDELized simulated galaxies



Summary

- Large Visual Classification effort within the team
 - 3.5/5 fields have been classified so far
 - 50,000 galaxies by the end
 - To be made public, GOODS-S imminently (Kartaltepe et al., in prep)
 - Lots of science based on these in progress!
- CANDELS and Galaxy Zoo – stay tuned!
- For more information about CANDELS, check out our blog at: <http://candels-collaboration.blogspot.com>
- Twitter: @CANDELS_team