

Galaxy Quenching and AGN Feedback

or, why the green valley is a red herring

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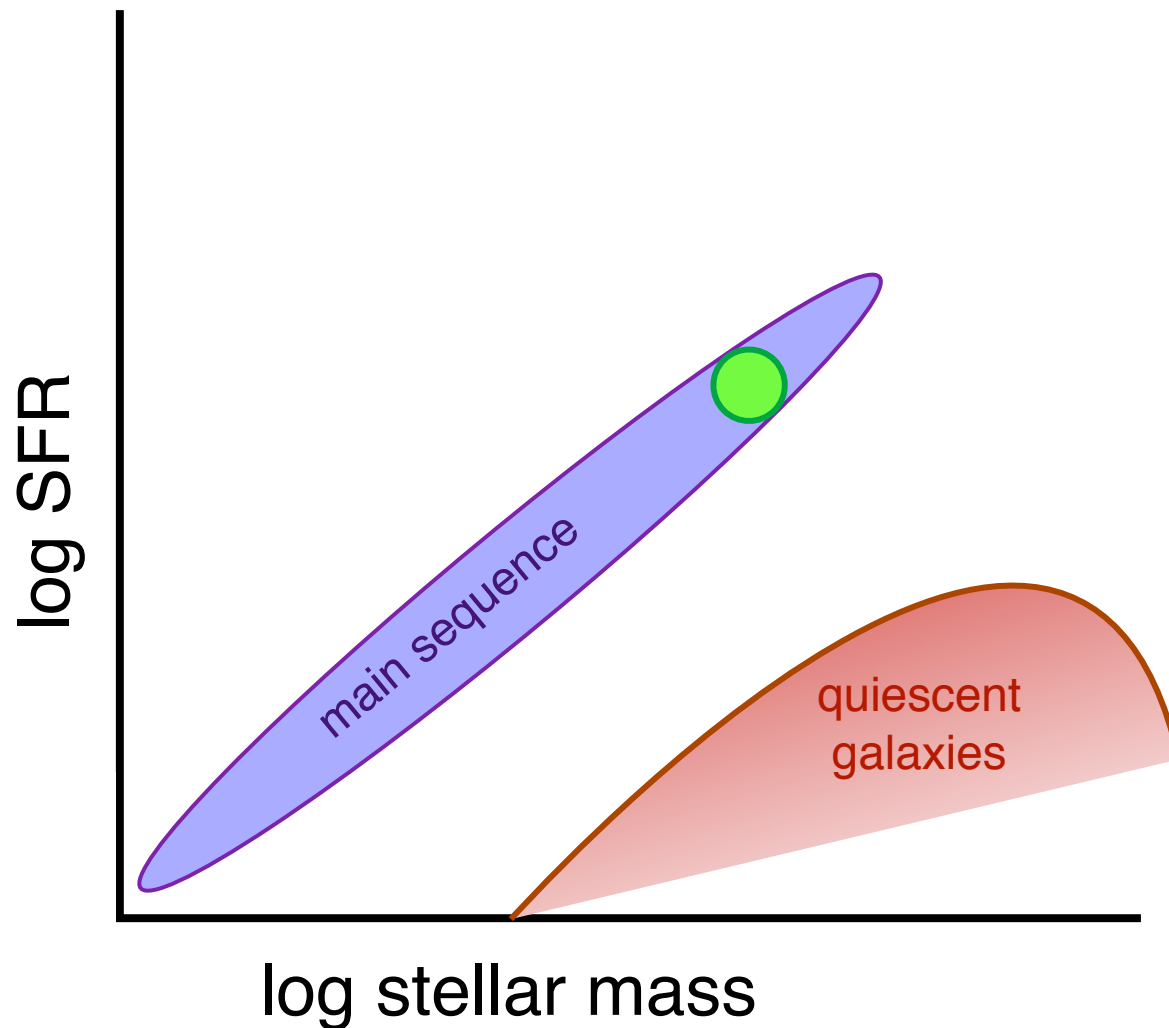
Eidgenössische Technische Hochschule Zürich
Swiss Federal Institute of Technology Zurich



@kevinschawinski

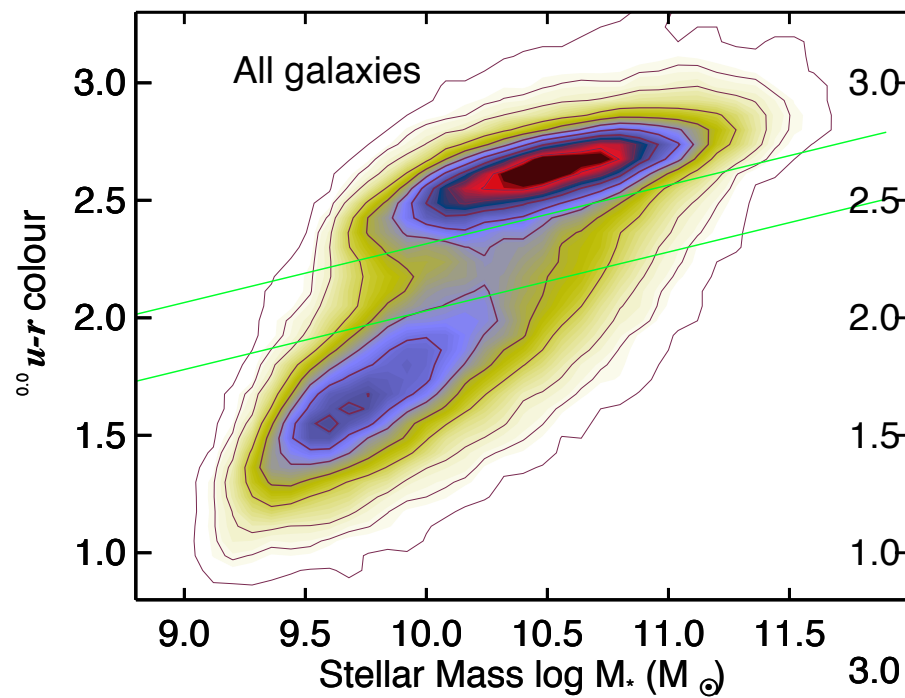
Context: the main sequence

(of star formation, don't throw things at me!)



Star forming galaxies are on main sequence: $M_{\text{stellar}} \sim \text{SFR}^\beta$, with small scatter. Scatter small, holds out to at least $z \sim 2$, perhaps higher.

See: Brinchmann+04; Elbaz+07,11, Noeske+07; Salim+07; Peng+10; Lee+12



Data

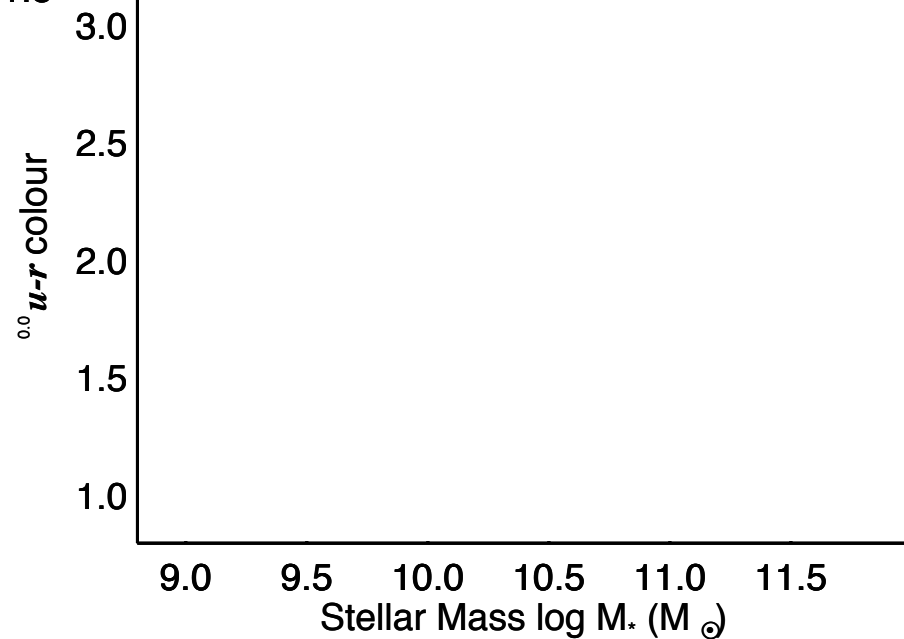
SDSS DR7

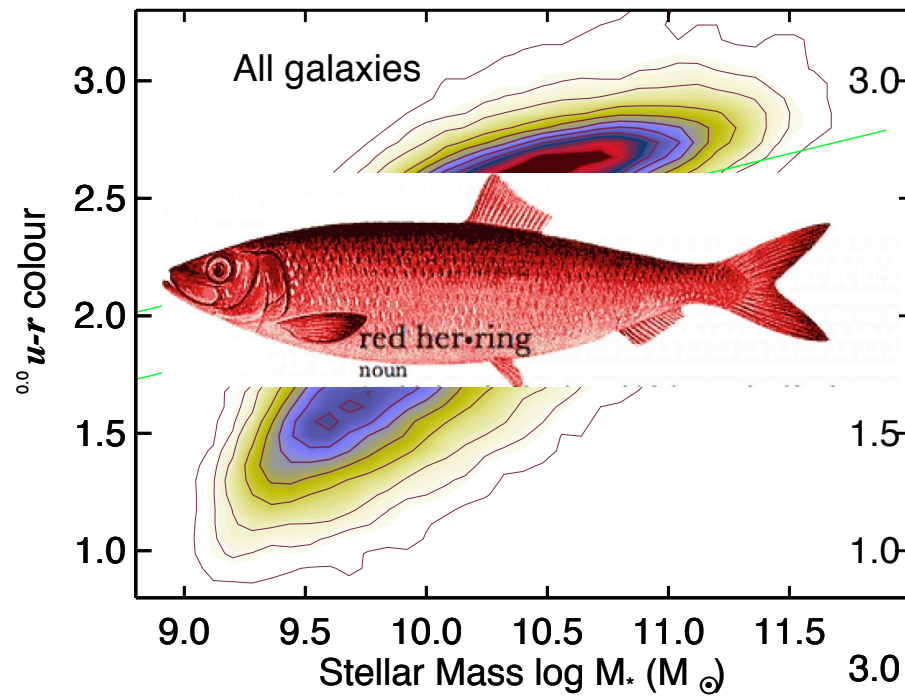
Galaxy Zoo 1 morphologies

GALEX GR6

MPA-JHU masses & SFRs

OSSY emission lines





Data

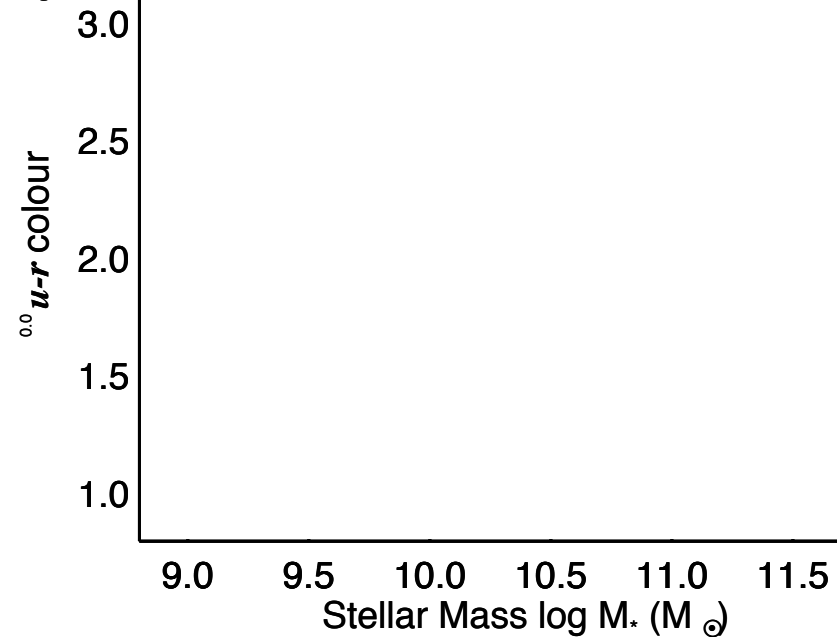
SDSS DR7

Galaxy Zoo 1 morphologies

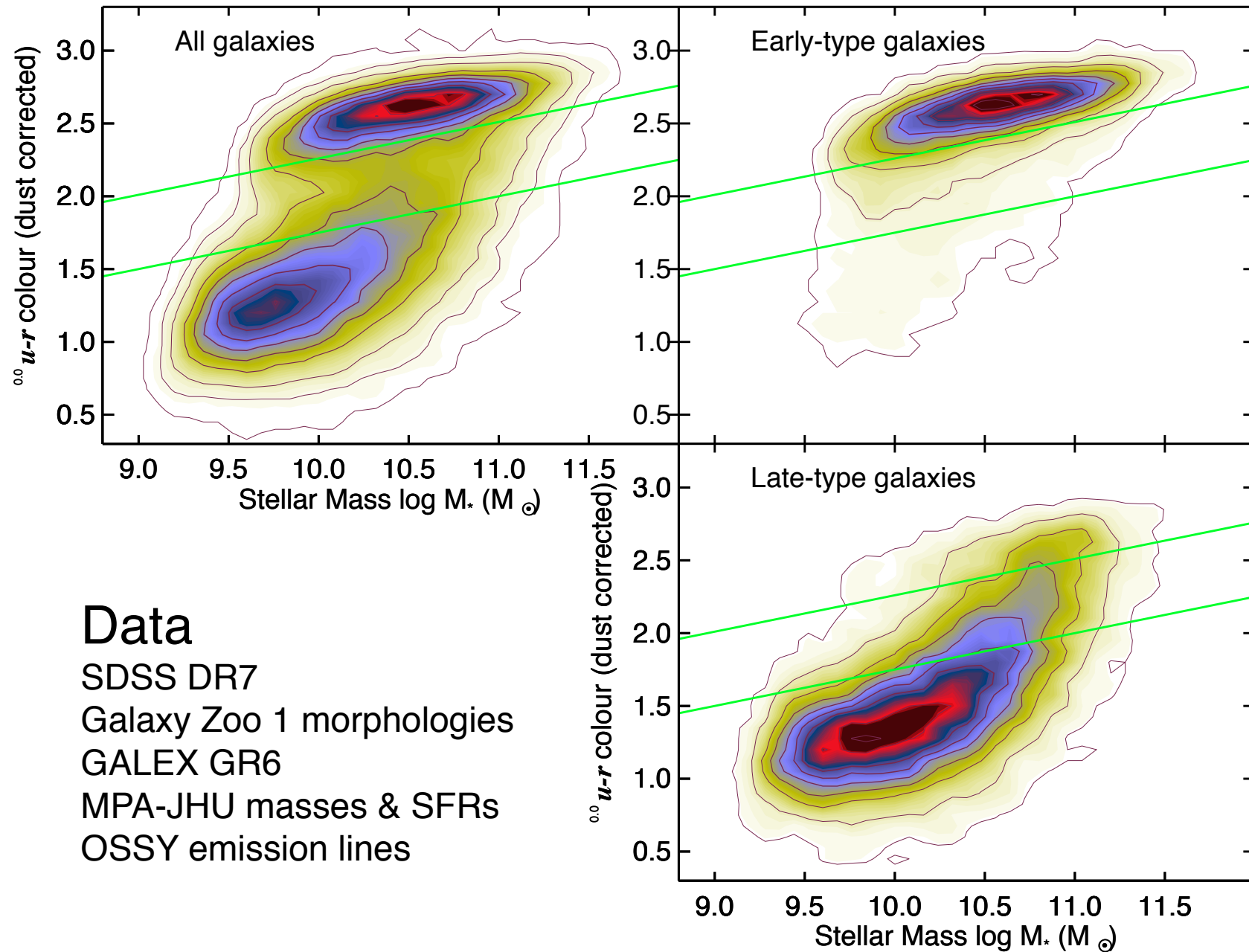
GALEX GR6

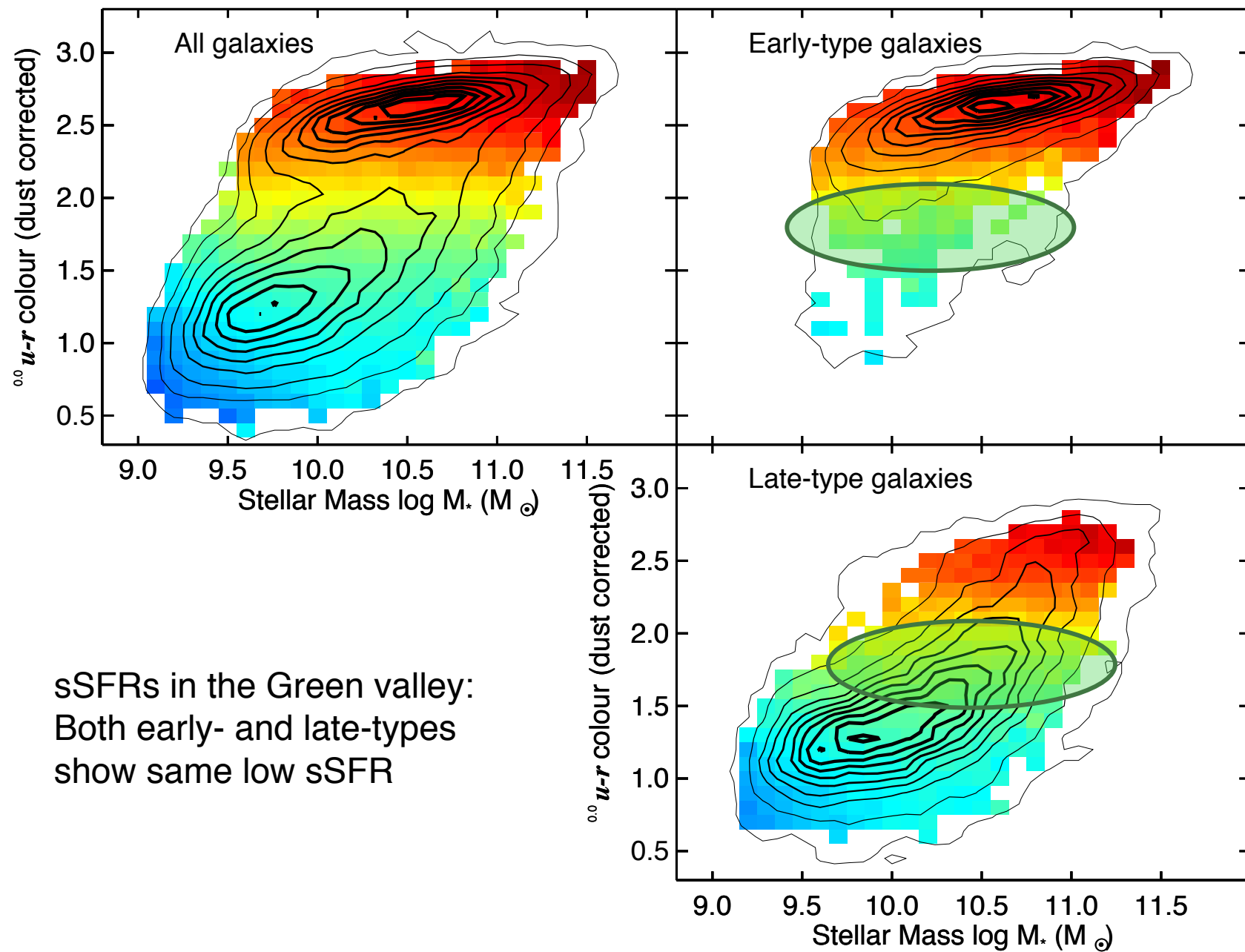
MPA-JHU masses & SFRs

OSSY emission lines



dust-corrected colours!



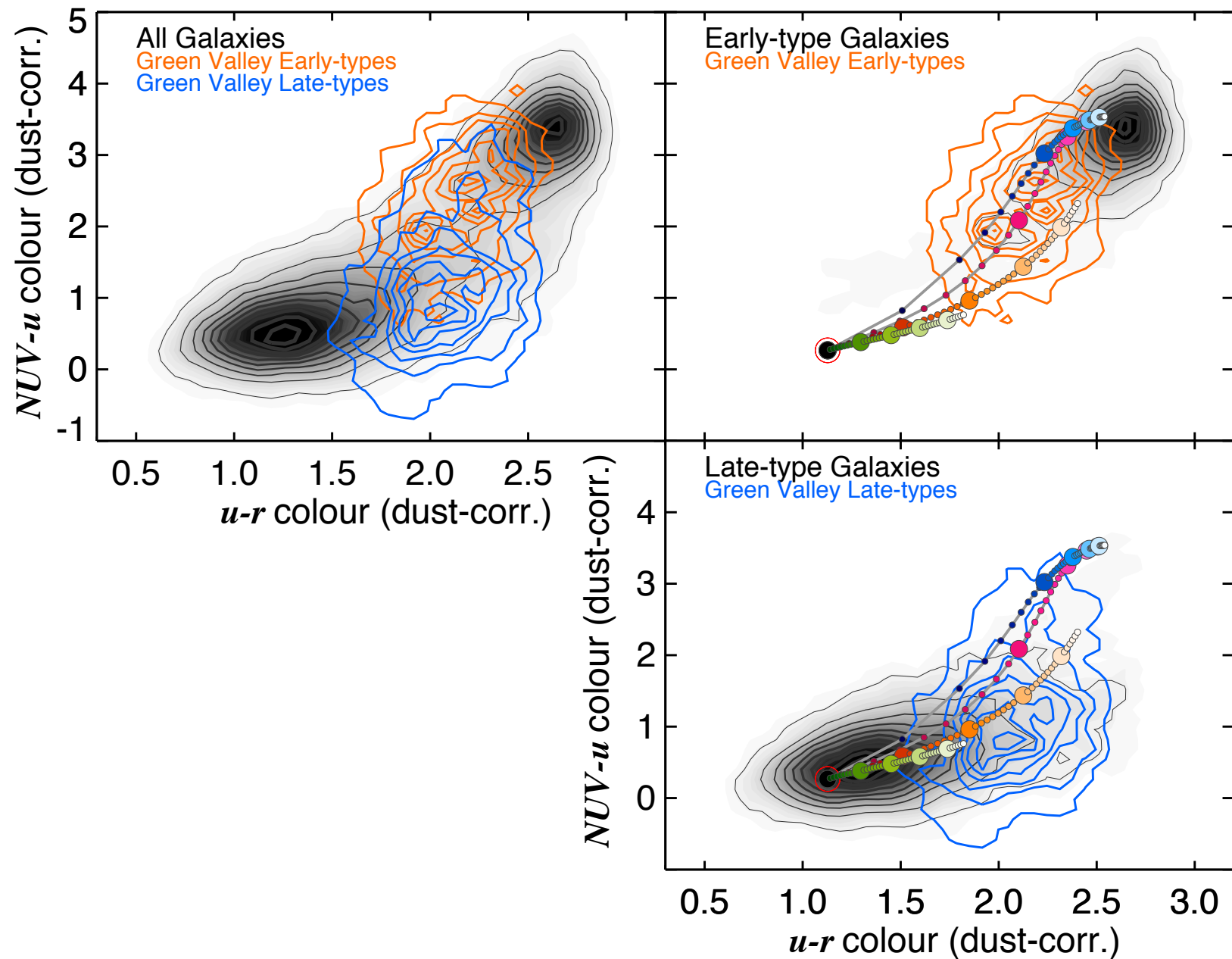


sSFRs in the Green valley:
Both early- and late-types
show same low sSFR

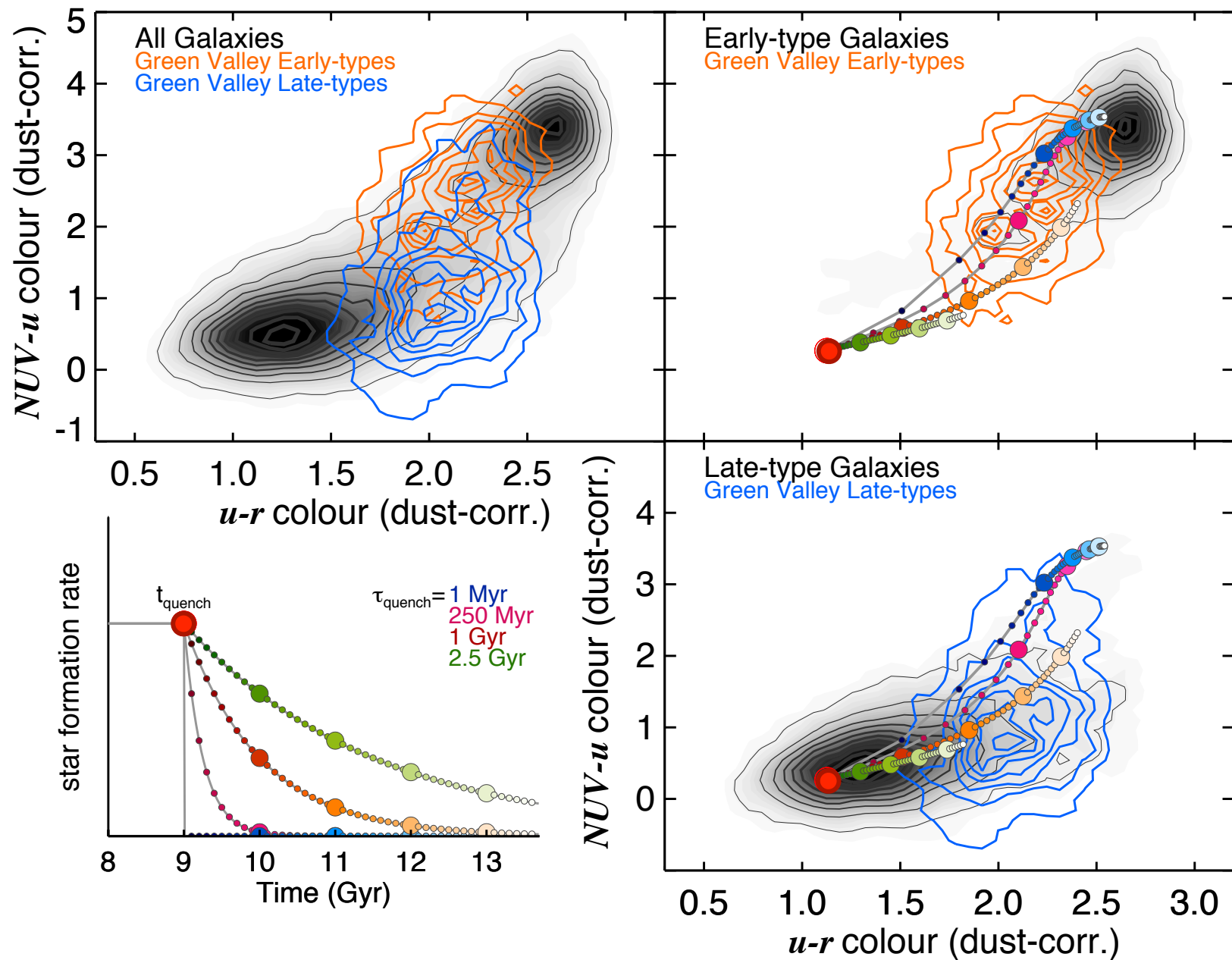
Measure* d/dt sSFR(t) of green valley galaxies

* really, broadly constrain for now

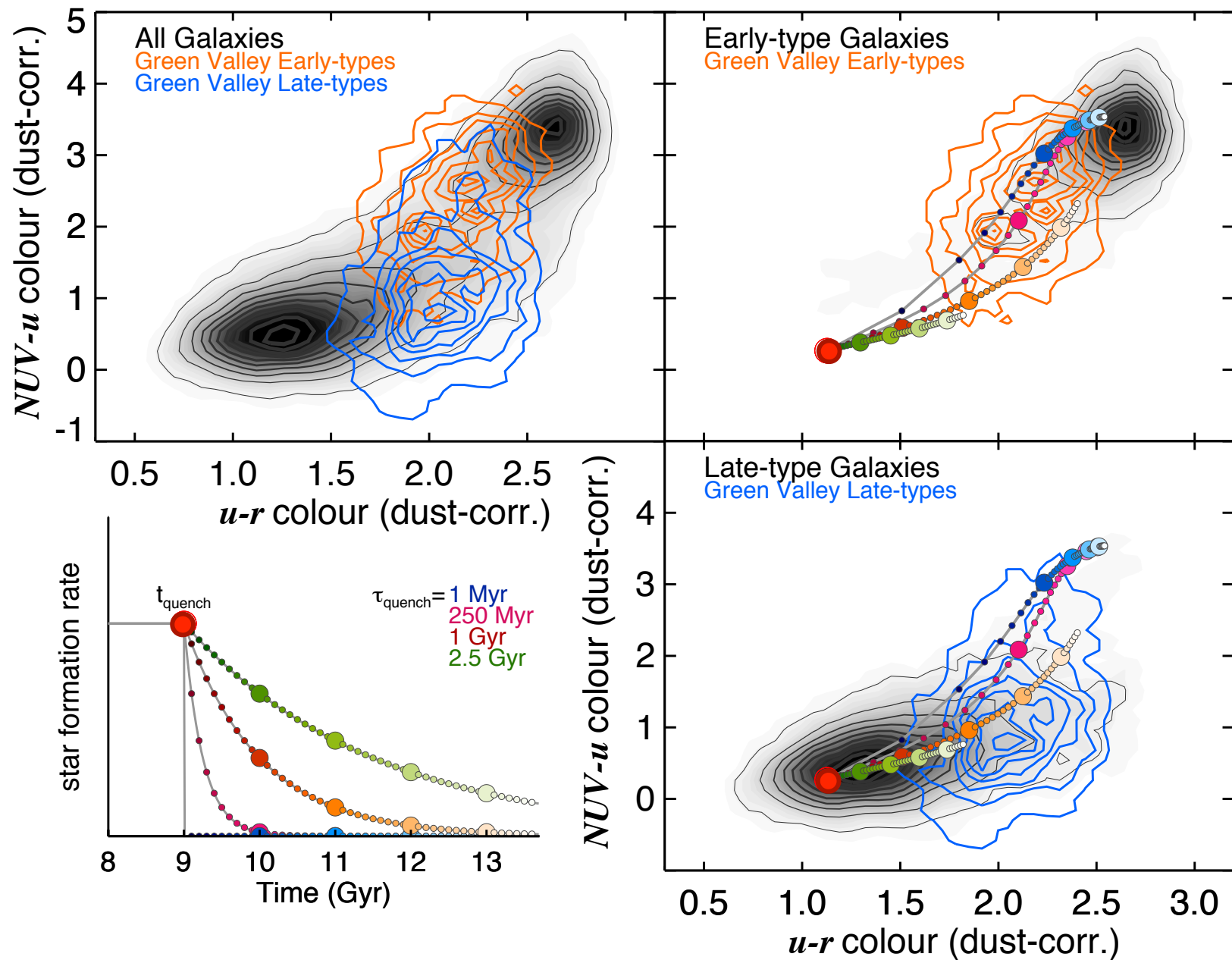
UV-optical colours constrain d/dt sSFR(t)



UV-optical colours constrain d/dt sSFR(t)

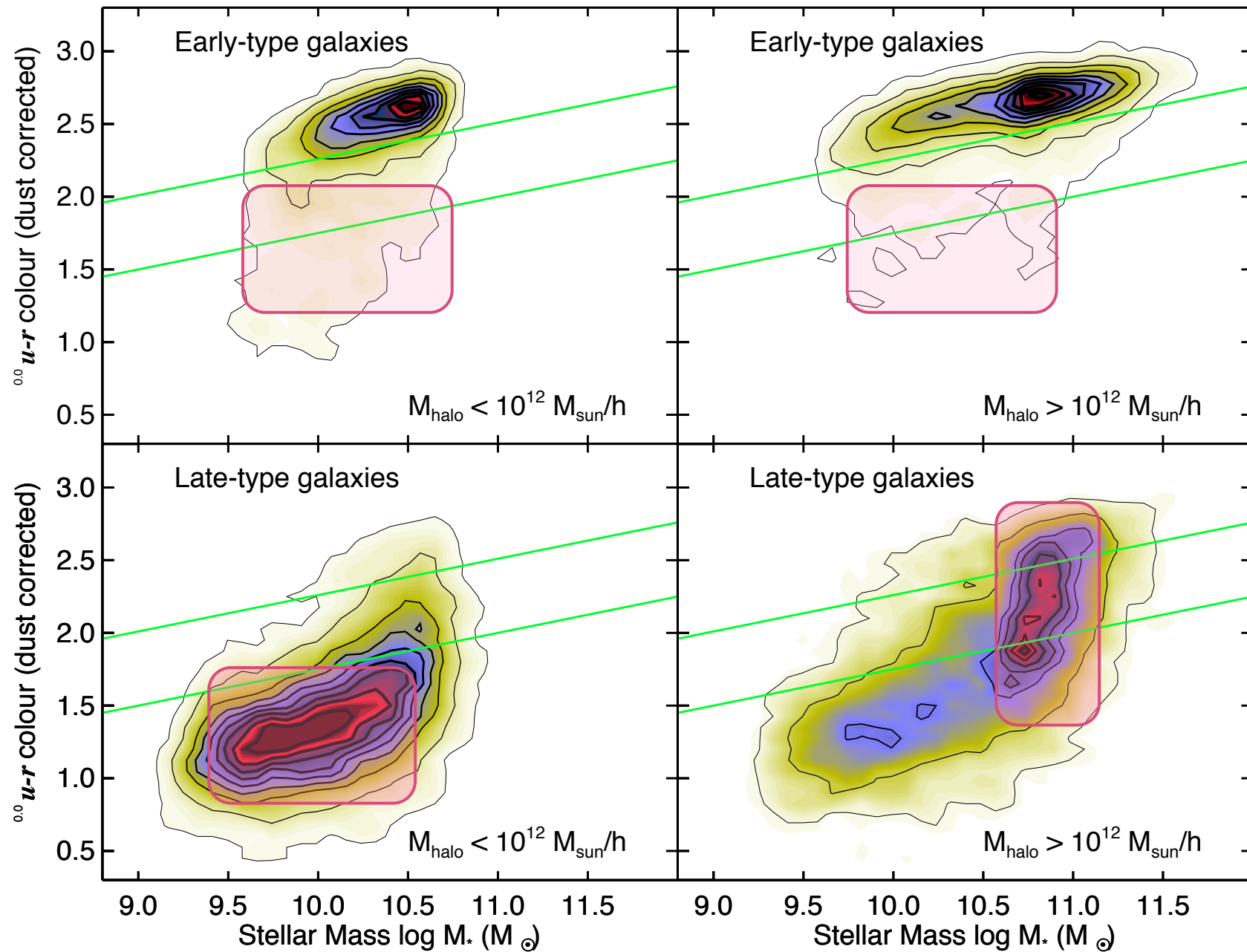


UV-optical colours constrain d/dt sSFR(t)



What about environment?
(quick detour)

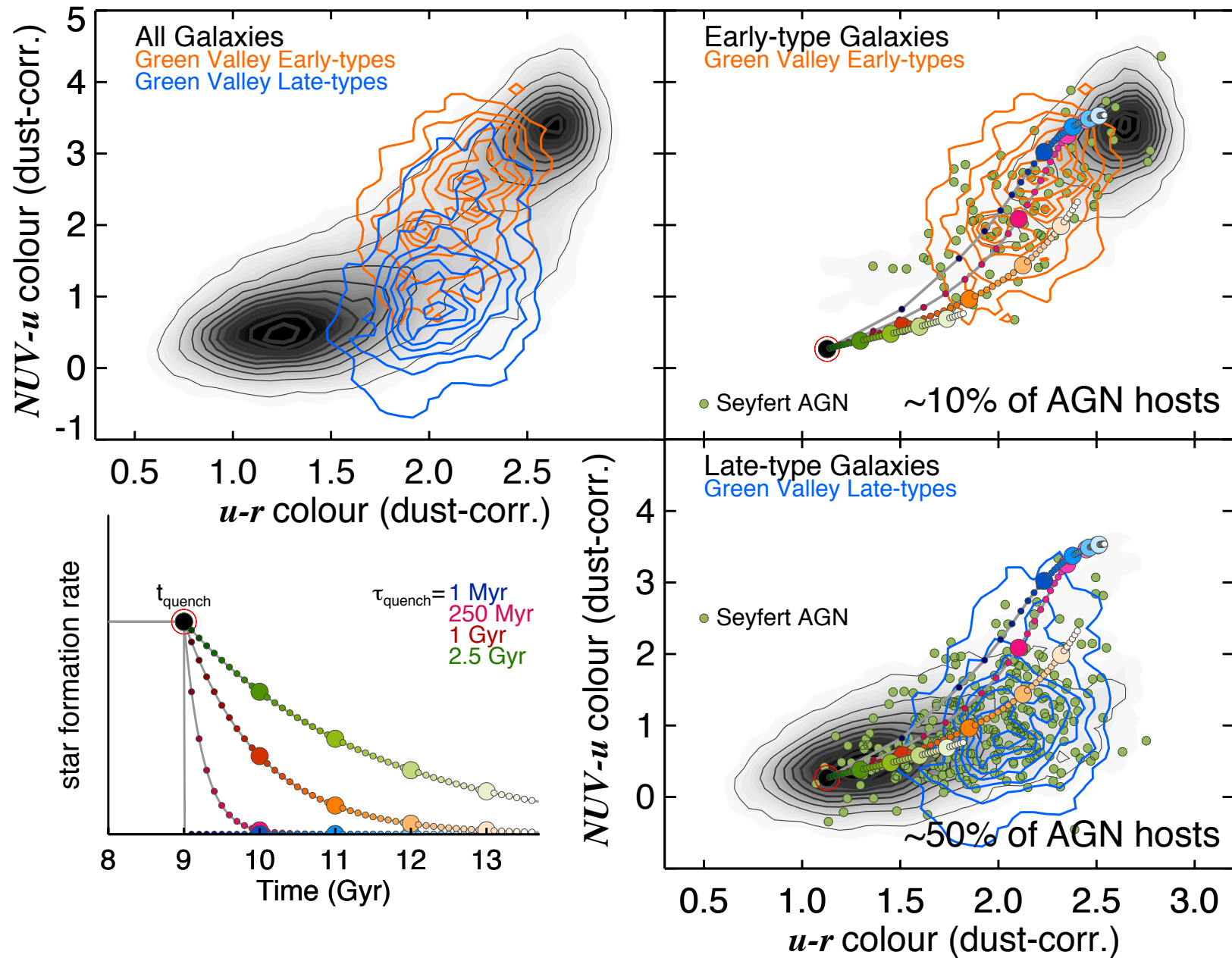
Environment



Halo masses from DR7 catalogue of Yang+

What about AGN?
Is there feedback?

Black Hole Growth



BPT AGN classification from Oh+12

$$\text{SFR} = \varepsilon M_{\text{gas}} / \tau_{\text{dyn}}$$

> $\text{SFR}(t) \sim \text{exponential}$

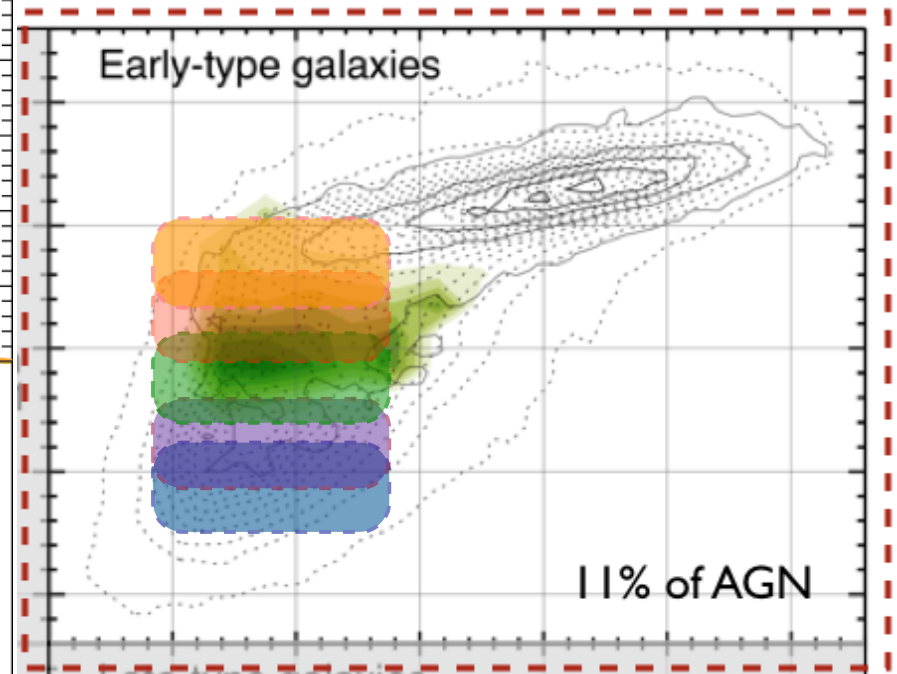
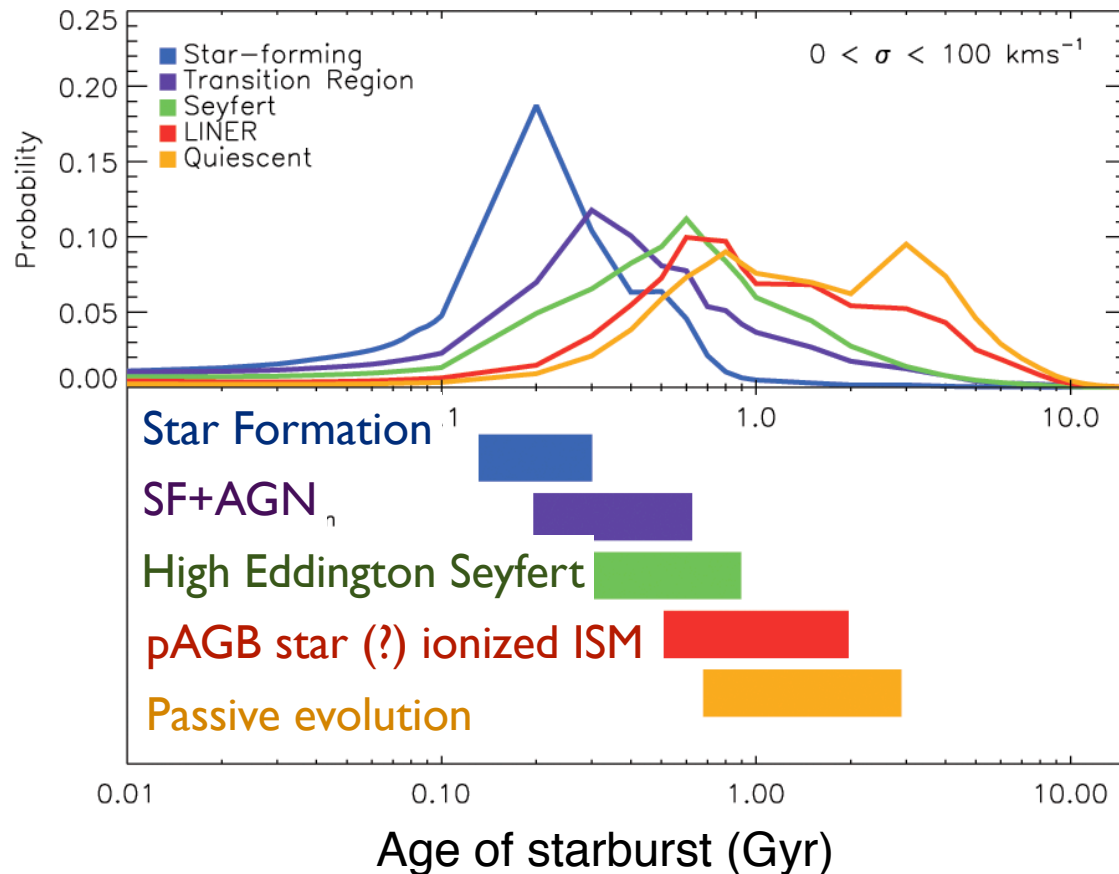
$$\varepsilon \sim 2\%$$

$$\tau_{\text{dyn}} \sim \text{few } 100 \text{ Myr}$$

(see Kaviraj+10!)

SDSS reveals an evolutionary sequence:

Low-mass early-types migrate to the red sequence

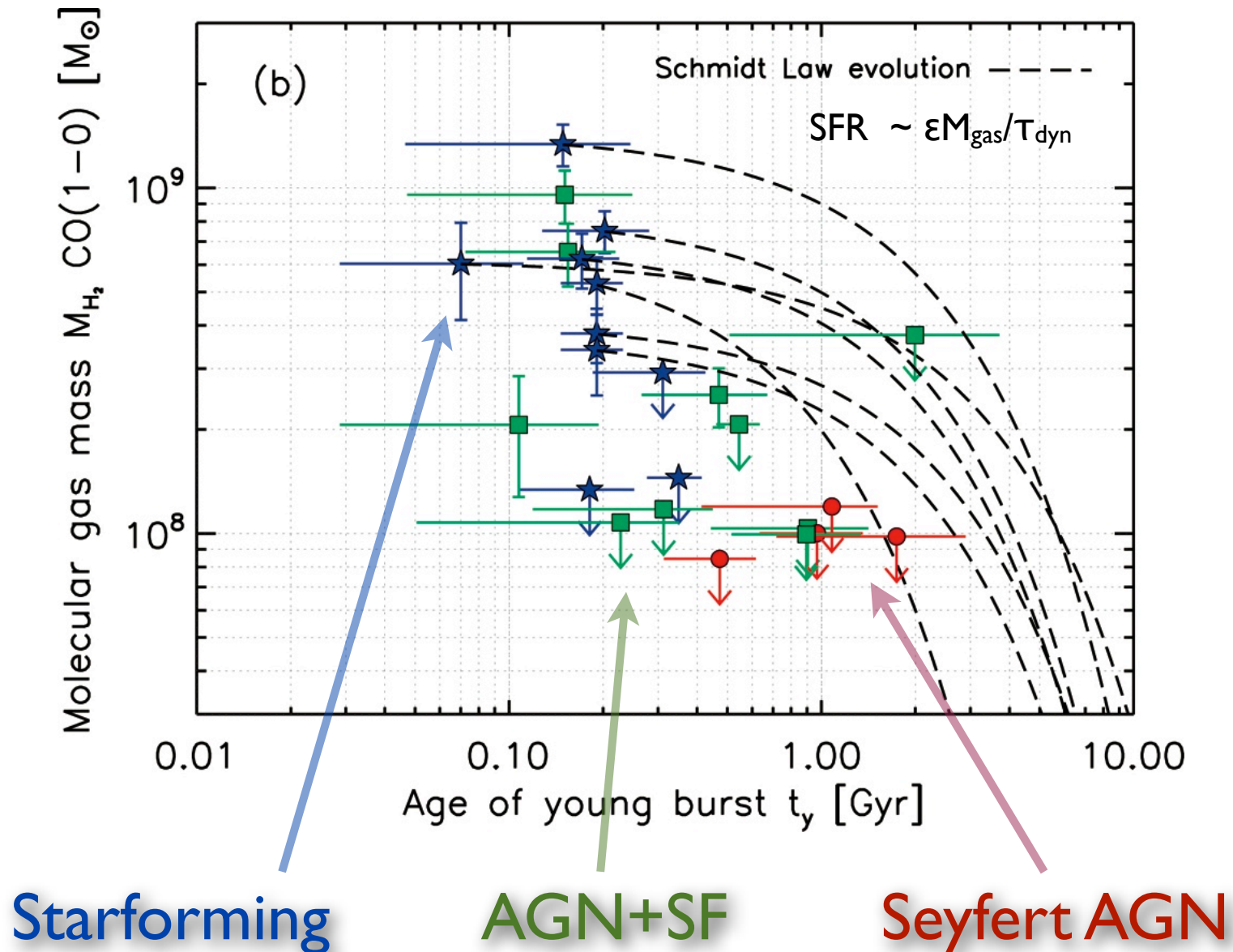


Schawinski+07

Input: GALEX+SDSS+2MASS photometry, SDSS spectroscopy (Lick indices)

Result: **1-10% of stellar mass** formed in most recent burst; leads to **rapid quenching** on ~100 Myr timescale and **migration to red sequence**

Deviations from the Schmidt Law: Clues to Feedback



Schawinski+09a, Kaviraj, Schawinski, Silk & Shabala 2010

Can we model* these two quenching pathways?

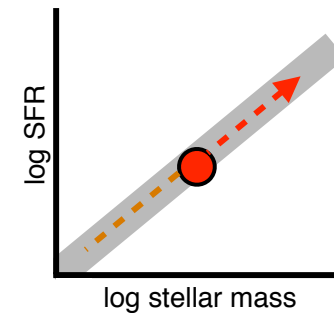
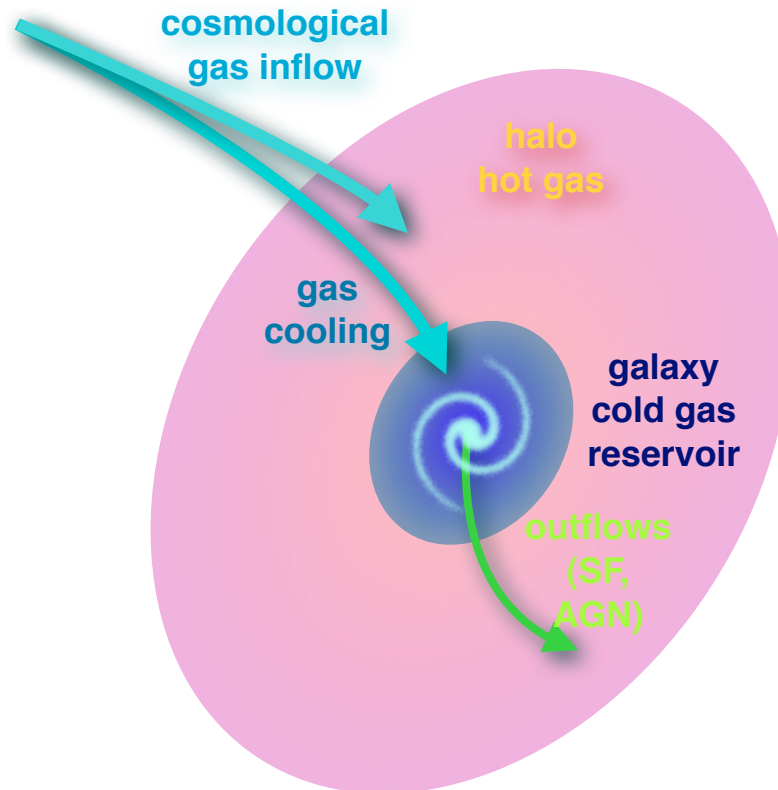
* really, very roughly

1. Late-type galaxies



Late-type galaxies quench star formation slowly

(1) prior to t_{quench}



Galaxy is on main sequence

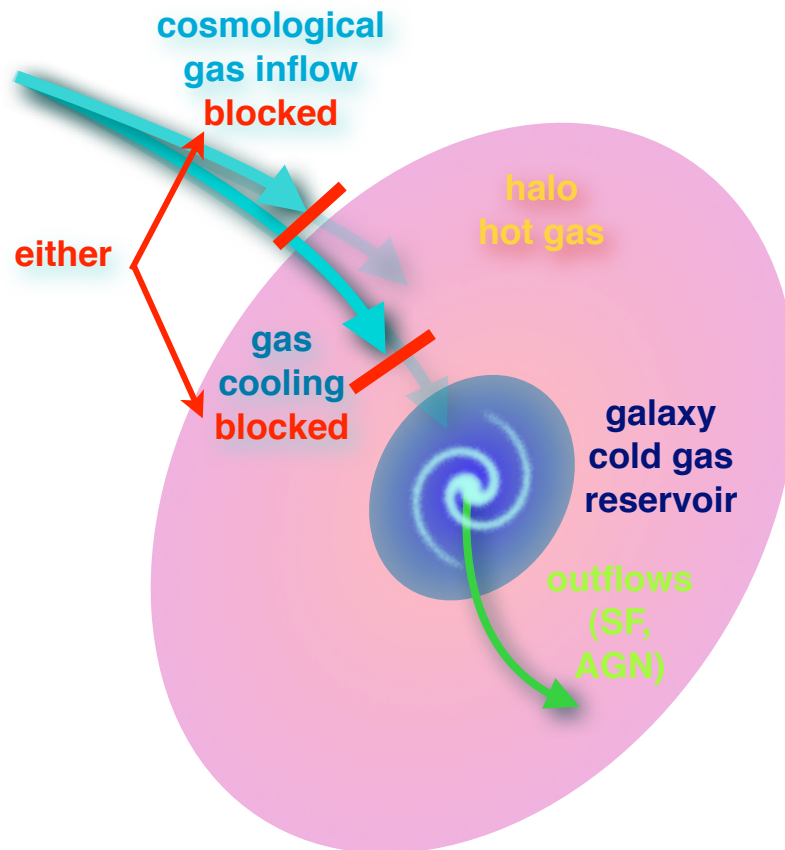
$$M_{\text{stellar}} \sim \text{SFR}^{\beta}$$

Inflows balance outflows:
system in quasi-equilibrium
(Bouche+10; Lilly+13)

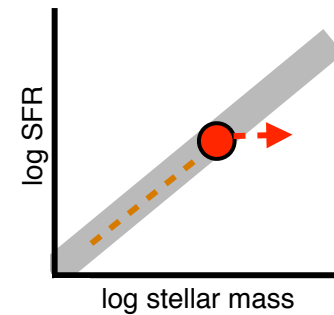
*could correspond to classical
environmental mechanisms (strangulation,
harassment, starvation).*

Late-type galaxies quench star formation slowly

(2) $t=t_{\text{quench}}$



could correspond to classical environmental mechanisms (strangulation, harassment, starvation).



Either cosmological inflows or cooling from halo are stopped, galaxy gas reservoir is now no longer replenished

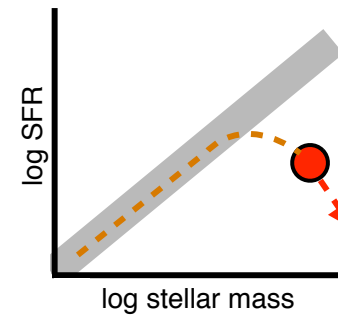
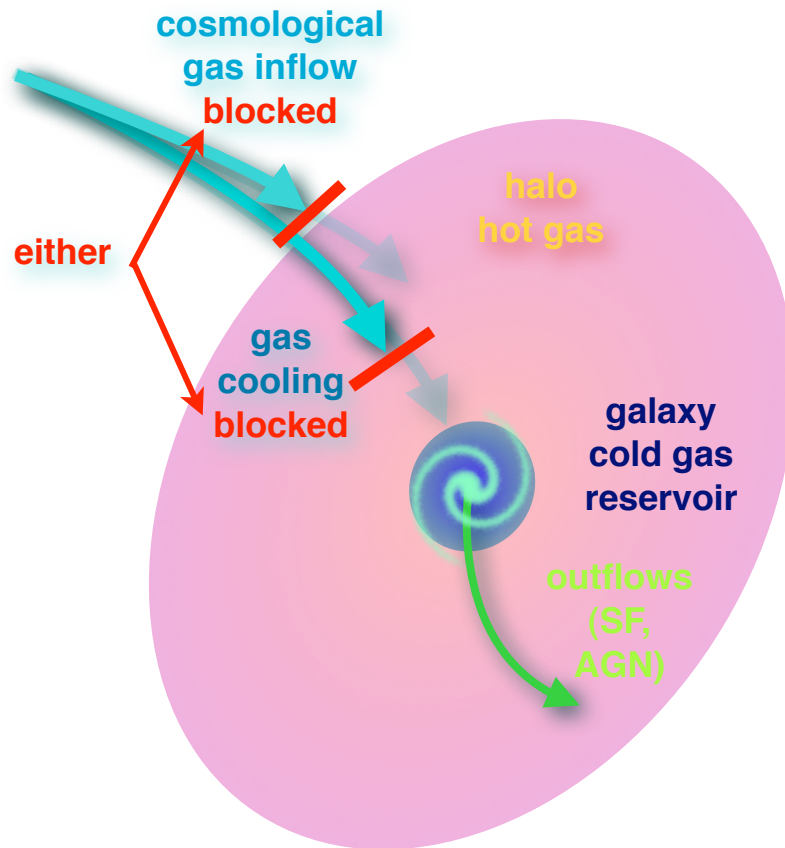
Galaxy leaves main sequence

$$\text{SFR} = \epsilon M_{\text{gas}} / \tau_{\text{dyn}}$$

$> \text{SFR}(t) \sim \text{exponential}$

Late-type galaxies quench star formation slowly

(3) several 100-1000 Myr past tquench



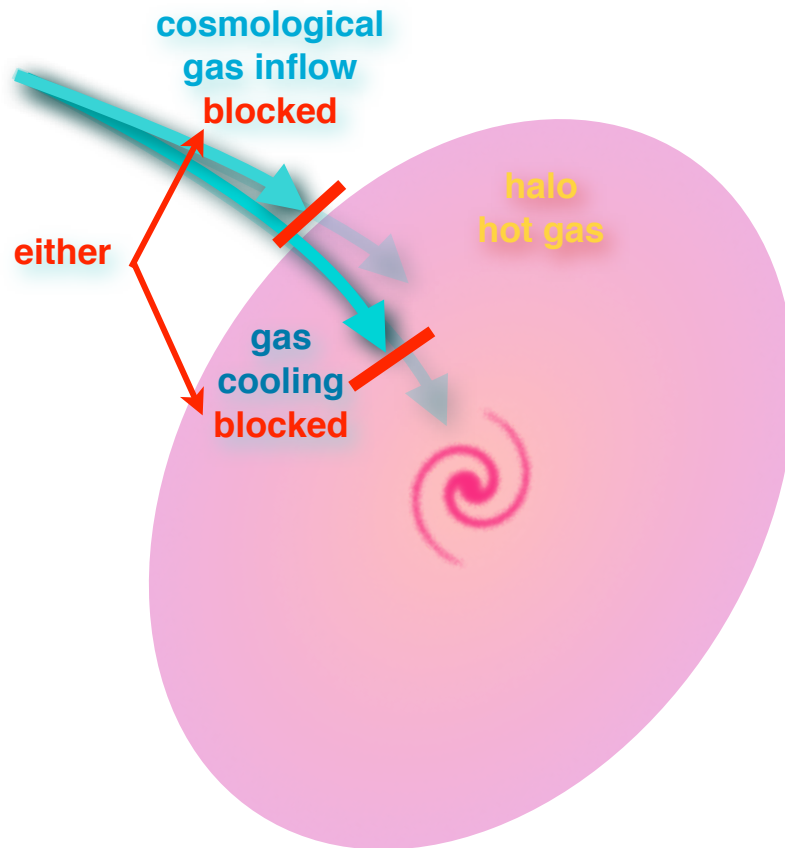
SFR goes into exponential decline, galaxy enters green valley

Gas reservoir is slowly used up

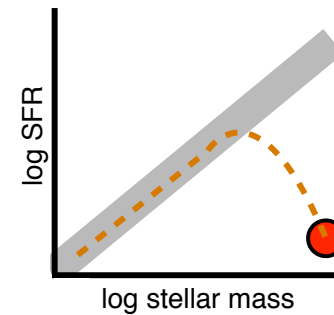
could correspond to classical environmental mechanisms (strangulation, harassment, starvation).

Late-type galaxies quench star formation slowly

(4) several Gyr past quench



could correspond to classical environmental mechanisms (strangulation, harassment, starvation).



Several Gyr later...

Passive, red spiral galaxy

No morphological transformation > “red spiral”

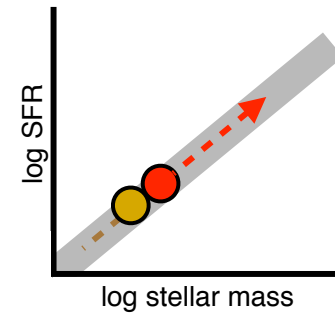
Very low SSFR

2. Early-type galaxies



Early-type galaxies quench star formation rapidly

(1) ~ 100 s Myr prior to t_{quench}

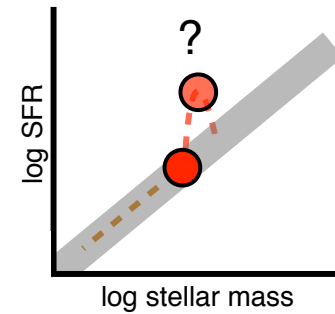
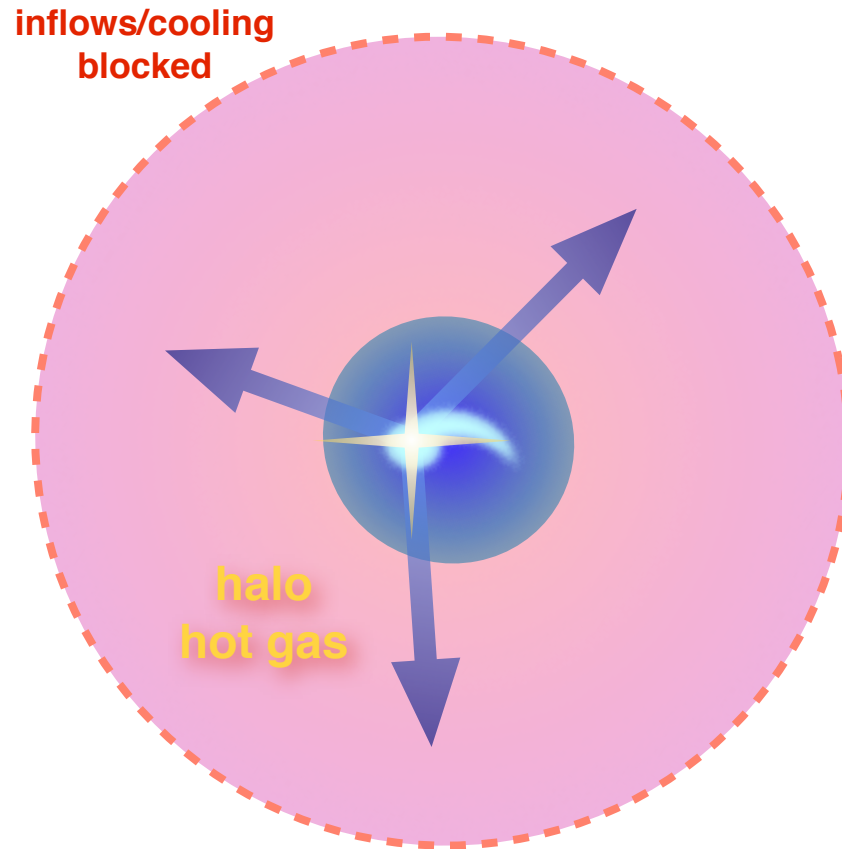


Two galaxies on main sequence
are about to merge

Both bring in gas reservoir and are
connected to cosmic gas inflow

Early-type galaxies quench star formation rapidly

(2) t_{quench}

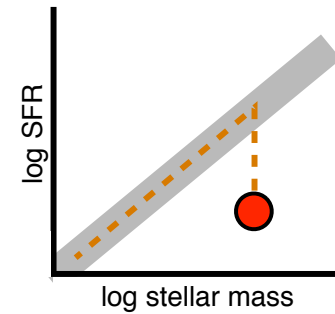
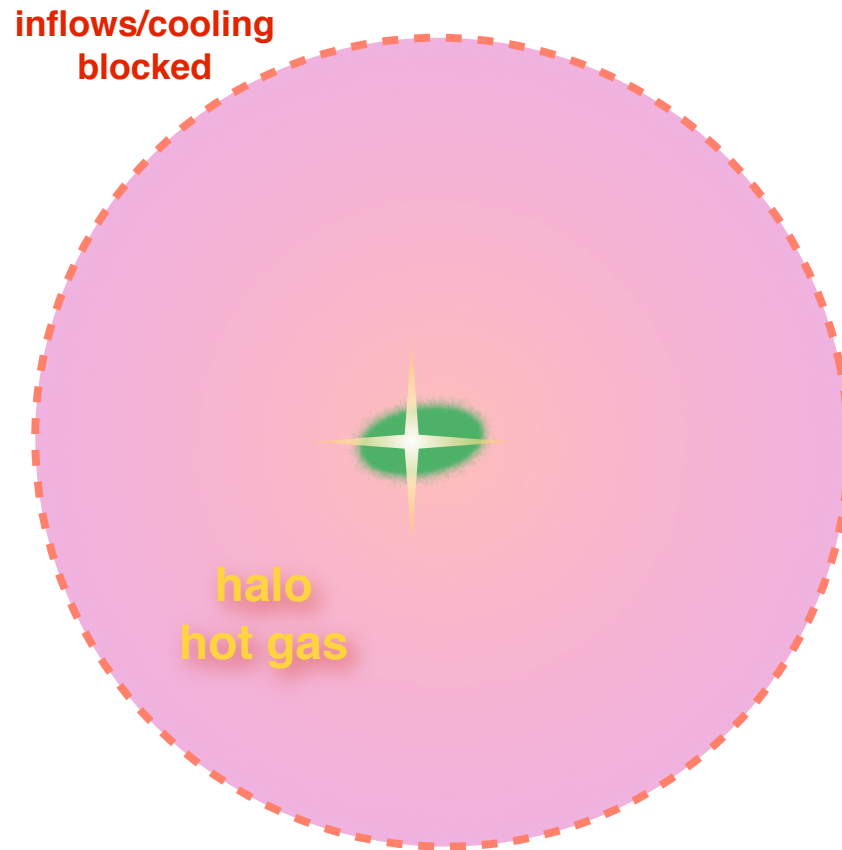


Merger may drive SFR above main sequence, or not. Morphology is transformed to spheroid

Some process drives outflows of gas to rapidly deplete the cold gas reservoir, perhaps AGN feedback?
 $\text{SFR} \neq \epsilon M_{\text{gas}} / \tau_{\text{dyn}}$
Further cosmological inflows and/or cooling stopped

Early-type galaxies quench star formation rapidly

(3) ~ 100 s Myr post t_{quench}



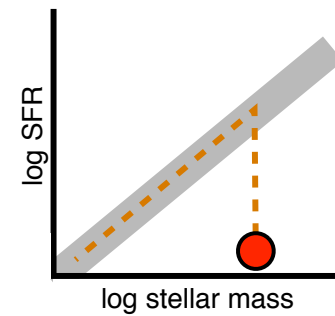
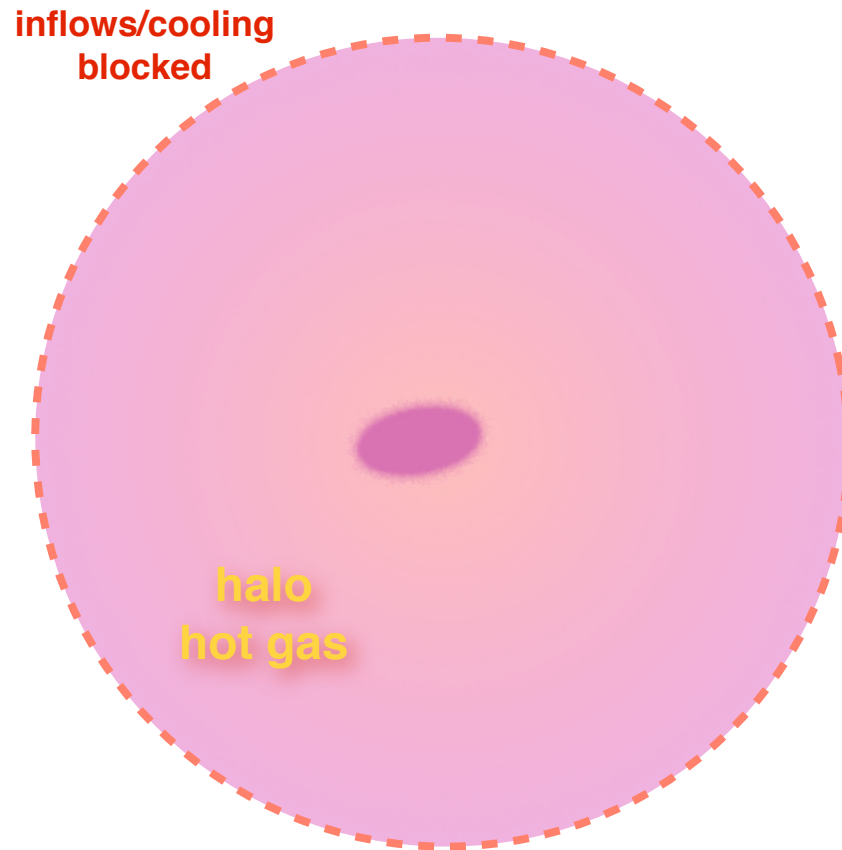
SFR drops rapidly to zero,
galaxy moves to green valley

Gas reservoir is destroyed,
further inflows not allowed to
replenish

AGN active

Early-type galaxies quench star formation rapidly

(4) $\sim 1\text{-}2$ Gyr post t_{quench}



With no gas reservoir and no further gas inflows, new early-type becomes passive red sequence galaxy.

Summary

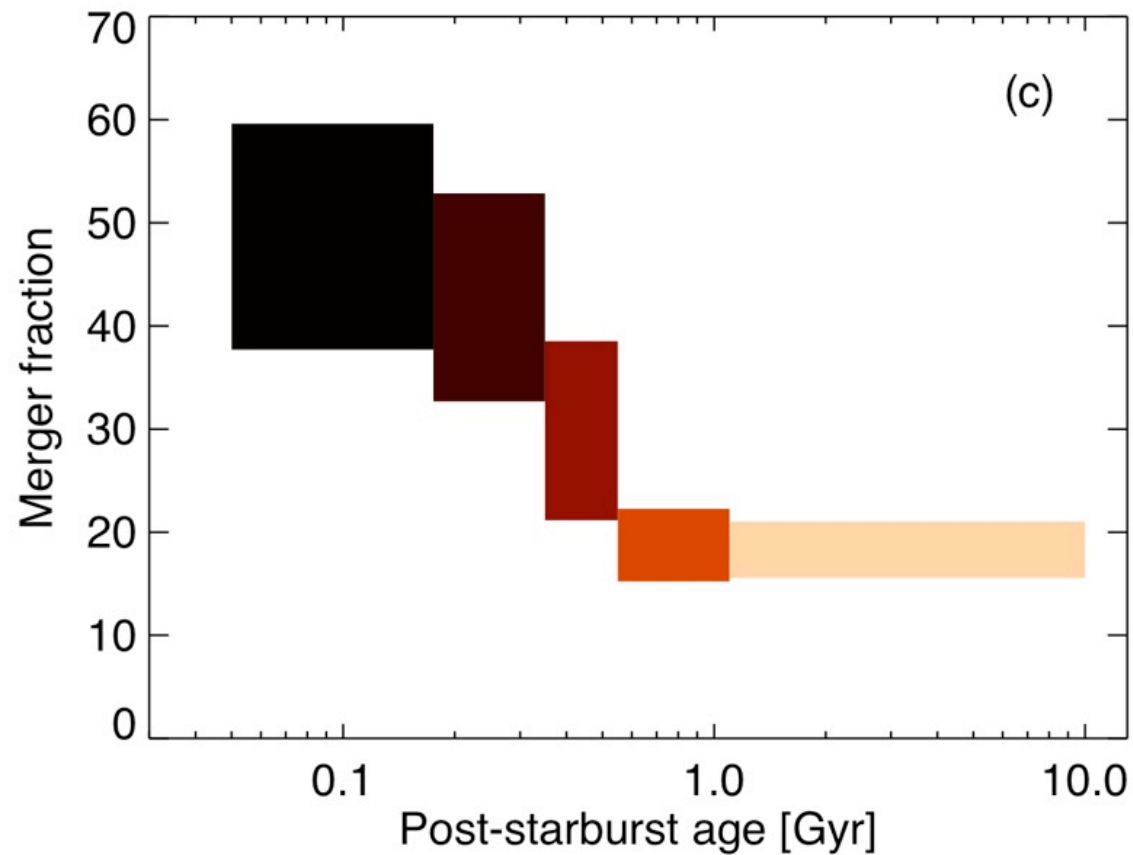
sSFR + UV/optical colours show **two different quenching pathways** though green valley associated with early- and late-type galaxies

Late-type galaxies **quench slowly, without changing morphology** (d/dt sSFR small) due to an external process that stops further inflow/cooling of gas; *could correspond to classical environmental mechanisms.*

Early-type galaxies **quench rapidly, during or after merger-driven change of morphology** (d/dt sSFR large) due to an internal process that destroys the gas reservoir, possibly *AGN feedback?*



Mergers trigger the migration from the blue cloud to the red sequence



Schawinski+10